# WELL-POSEDNESS OF REYNOLDS AVERAGED EQUATIONS FOR COMPRESSIBLE FLUIDS WITH A VANISHING PRESSURE 

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#### Abstract

We show that the Reynolds averaged equations for compressible fluids (neglecting third order correlations) are well-posed in H s when the pressure vanishes in dimensions $d=2$ and 3 . In order to do this, we show that the system is Friedrichs-symmetrizable. This model belongs to the class of non-conservative hyperbolic systems. Hence the usual symmetrisation method for conservation laws can not be used here.


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We show that the Reynolds averaged equations for compressible fluids (neglecting third order correlations) are well-posed in $H^{s}$ when the pressure vanishes in dimensions $d=2$ and 3 . In order to do this, we show that the system is Friedrichs-symmetrizable. This model belongs to the class of non-conservative hyperbolic systems. Hence the usual symmetrisation method for conservation laws can not be used here.


## 1. Introduction and main results

We study the Reynolds averaged equations for compressible fluids, where thirdorder correlations are neglected. This system can be written in Eulerian coordinates as

$$
\left\{\begin{array}{l}
\partial_{t} \rho+\operatorname{div}(\rho v)=0  \tag{1.1a}\\
\partial_{t} v+(v \cdot \nabla v)+\frac{1}{\rho}\left(\nabla p+\operatorname{div}(\rho P)^{T}\right)=0 \\
\partial_{t} P+(v \cdot \nabla) P+\frac{\partial v}{\partial x} P+P\left(\frac{\partial v}{\partial x}\right)^{T}=0
\end{array}\right.
$$

The variables are the averaged density $\rho>0$, the averaged velocity $v \in \mathbb{R}^{d}$ $(d=2$ or 3$)$, and $P$ is the Reynolds stress tensor, $P \in S_{d}^{++}(\mathbb{R})$. The function $p$ is the pressure of the fluid and is a function of the density $\rho$, through an equation of state $p=p(\rho)$. The map $\rho \mapsto p(\rho)$ is supposed to be of class $C^{1}$ and nondecreasing. Typical pressure laws are of the form $p(\rho)=a \rho^{\gamma}$, with $a>0$ and $\gamma>0$ two constants.

The tensor $P$ is a classical Reynolds tensor appearing in the Reynolds averaging of turbulent flows for barotropic compressible fluids (see [14, [18, [12]). It also appears in the description of free surface shear flows, where the averaging operator is the depth averaging (cf. Annexe C, A7 in [15] for a derivation of the model). In that latter case, the density $\rho$ must be replaced by the water depth, often denoted $h$. The pressure is then given by $p(h)=g h^{2} / 2$ (cf. [15] for instance).

System (1.1) is hyperbolic whenever $p^{\prime}(\rho) \geq 0$ and $P$ is definite positive, as it has been proved in [15]. Equation 1.1a) shows that the density $\rho$ is conserved. The conservation of momentum $\rho v$ also holds: 1.1 b rewrites as

$$
\begin{equation*}
\partial_{t}(\rho v)+\operatorname{div}\left(\rho v \otimes v+p I_{d}+\rho P\right)^{T}=0 . \tag{1.2}
\end{equation*}
$$

[^0]One can also deduce from (1.1) the conservation of energy:

$$
\begin{equation*}
\partial_{t} e+\operatorname{div}(e v+p v+\rho P v)=0, \text { with } e:=\frac{1}{2} \rho|v|^{2}+\rho \mathcal{E}(\rho)+\frac{1}{2} \operatorname{Tr}(\rho P) \tag{1.3}
\end{equation*}
$$

The map $\rho \mapsto \rho \mathcal{E}(\rho)$ is called the volumic internal energy, and is linked to the pressure via the relation $p(\rho)=\rho^{2} \mathcal{E}^{\prime}(\rho)$. The term $\rho|v|^{2} / 2$ is the volumic kinetic energy of the fluid, and the term $\operatorname{Tr}(\rho P) / 2$ is the energy associated to the tensor $P$.

In [6], it was shown that system (1.1) admits a variational formulation, as it is often the case in physics when the energy of a system is conserved. Define the Lagrangian density

$$
\begin{equation*}
\mathcal{L}(\rho, v, P):=\frac{1}{2} \rho|v|^{2}-\rho \mathcal{E}(\rho)-\frac{1}{2} \operatorname{Tr}(\rho P), \tag{1.4}
\end{equation*}
$$

and the corresponding action

$$
\begin{equation*}
\mathcal{A}:=\int_{t_{0}}^{t_{1}} \int_{\mathbb{R}^{d}} \mathcal{L} \mathrm{~d} x \mathrm{~d} t \tag{1.5}
\end{equation*}
$$

Then one can show that 1.1 b is the Euler-Lagrange equation given by the stationnary action principle applied to the action 1.5, under the two constraints 1.1a and (1.1c).

The tensor $P$ admits an additional conservation law, sometimes called conservation of enstrophy, that is a consequence of 1.1 c and can be written

$$
\begin{equation*}
\partial_{t}\left(\frac{\operatorname{det} P}{\rho^{2}}\right)+v \cdot \nabla\left(\frac{\operatorname{det} P}{\rho^{2}}\right)=0 \tag{1.6}
\end{equation*}
$$

Note that equation $\sqrt{1.1 \mathrm{c}}$ implies that the symmetry of $P$ is conserved by the evolution. Equation (1.6) then implies that, if $P(t) \in S_{d}^{++}(\mathbb{R})$ for some instant $t$, then this property is true for all times.

It has been proved in [7] that system (1.1) does not admit any further conservation law. Thus, in dimension $d=2$ or 3 , system (1.1) is not conservative. Hence the usual symmetrisation method of Godunov (cf. [8]) and Lax and Friedrichs (cf. [5]) for hyperbolic systems of conservation laws (see for instance [17], pages 83-84) can not be used here.

However, one can show that system (1.1) is Friedrichs-symmetrizable when the pressure vanishes. More precisely, we state the following proposition:

Proposition 1. Let $d=2$ or 3 . Suppose that $\rho$ takes values in $\mathbb{R}_{+}^{*}$ and $P$ takes values in $S_{d}^{++}(\mathbb{R})$. Then the two following properties are equivalent:
(1) System 1.1 is Friedrichs-symmetrizable
(2) The pressure $p$ is constant: $p^{\prime}(\rho)=0$, or the tensor $P$ is a scalar matrix, i.e. there exists $\lambda=\lambda(t, x) \in \mathbb{R}$ such that $P=\lambda I_{d}$.

Let us make few comments about this proposition:

- The tensor $P=\lambda I_{d}$ is a solution of system 1.1 only for trivial velocities; hence, for applications of this result, the case $p^{\prime}=0$ seems more interesting.
- This property holds in the variables $(\rho, v, P)$. It could be possible that system (1.1), written in different variables, appears to be symmetrizable even when $p^{\prime} \neq 0$.
- When $d=1$, system (1.1) is symmetrizable, even when $p^{\prime} \neq 0$. In fact, one can prove that one-dimensional hyperbolic systems are always Friedrichssymmetrizable (cf. [13]). In higher dimension $d \geq 2$, this does not hold anymore.
As a consequence of Proposition 1, we have the following result regarding the well-posedness of system 1.1):

Theorem 1. Let $d=2$ or $3, s>1+d / 2$ and $\mathcal{U}:=\mathbb{R}_{+}^{*} \times \mathbb{R}^{d} \times S_{d}^{++}(\mathbb{R})$. Let $\bar{Y}:=$ $(\bar{\rho}, \bar{u}, \bar{P}) \in \mathcal{U}$ and $Y_{0}$ taking values in $\mathcal{U}$ such that $Y_{0}-\bar{Y} \in H^{s}\left(\mathbb{R}^{d}\right)$. We consider the Cauchy problem associated to (1.1) with initial data $Y_{0}$. There exists $T>0$ such that (1.1) with $p^{\prime}=0$ has a unique classical solution $Y(t)$ in $\mathcal{C}^{1}\left([0, T] \times \mathbb{R}^{d}\right)$ with values in $\mathcal{U}$ achieving the initial data $Y(0)=Y_{0}$. Furthermore, $Y-\bar{Y}$ belongs to $\mathcal{C}\left([0, T], H^{s}\left(\mathbb{R}^{d}\right)\right) \cap \mathcal{C}^{1}\left([0, T], H^{s-1}\left(\mathbb{R}^{d}\right)\right)$.

Theorem 1 is a consequence of Proposition 1 and of the theory of well-posedness of Kato for quasilinear evolution equations (cf. [11]). For a detailed proof of this result, see for instance Sect. 10 in [1].

System (1.1) has been used over the last years in the modeling of turbulent flows, including for numerical simulations (see [16, [10, [7, [9] ). However, the well-posedness of (1.1) in dimension $d \geq 2$ is still uncertain today. Theorem 1 states an answer to this question in the case of a vanishing pressure. One could also obtain system 1.1 with $p^{\prime}=0$ when modeling a fluid for which the pressure gradient $\nabla p$ is negligible compared to the turbulence of the fluid $\operatorname{div}(\rho P)$ in 1.1 b$)$.

The hypothesis $p^{\prime}=0$ can also be found in the litterature, in a model called "pressureless gas dynamics". The pressureless Euler equations are obtained from the Euler equations with a pressure assumed to be 0 . The pressureless gas dynamics are used to model astrophysical systems (see 19] for instance) and is only weakly hyperbolic. As a consequence, phenomena like creation of vacuum or high concentrations (delta shocks) can occur (see for instance [3], [4], [2]). System (1.1) with $p^{\prime}=0$ can thus be seen as a hyperbolic version of the pressureless gas model. In this sense, the Reynolds tensor $P$ brings more regularity to the model.

## 2. Proof of the proposition

Notations. We denote $\partial_{i}:=\partial / \partial x_{i}$ the partial derivative of the variable $x_{i}$, for $1 \leq i \leq d$. If $f: \mathbb{R}^{d} \rightarrow \mathbb{R}$ is a scalar function, we denote by $\nabla f \in \mathbb{R}^{d}$ the gradient of $f$, i.e. the vector field of components $\partial_{i} f, 1 \leq i \leq d$.

If $Z=\left(Z_{1}, \ldots, Z_{n}\right): \mathbb{R}^{d} \rightarrow \mathbb{R}^{d}$ is a vector field, the divergence of $Z$ is the scalar function defined by

$$
\operatorname{div}(Z):=\partial_{1} Z_{1}+\cdots+\partial_{n} Z_{n}
$$

We also denote by $\partial Z / \partial x$ the Jacobian matrix of $Z$, i.e. the matrix of coefficients $(\partial Z / \partial x)_{i, j}=\partial Z_{i} / \partial x_{j}$, for $1 \leq i, j \leq d$.

If $Z=\left(Z_{i}\right)_{1 \leq i \leq d}$ and $Z^{\prime}=\left(Z_{i}^{\prime}\right)_{1 \leq i \leq d}$ are two vector fields, we denote $Z \otimes Z^{\prime}$ the second order tensor defined by $Z \otimes Z^{\prime}:=Z\left(Z^{\prime}\right)^{T}$, i.e. the matrix of coefficients $\left(Z \otimes Z^{\prime}\right)_{i, j}=Z_{i} Z_{j}^{\prime}$, for $1 \leq i, j \leq d$.

If $A: \mathbb{R}^{d} \rightarrow M_{d}(\mathbb{R})$ is a second order tensor, we defined the divergence of $A$ as the line vector of $\mathbb{R}^{d}$ whose $i$-th component is given by the divergence of the $i$-th column of $A$.

For any positive integer $d$, we denote $I_{d} \in M_{d}(\mathbb{R})$ the identity matrix of size d. We denote $S_{d}^{++}(\mathbb{R})$ the set of symmetric definite positive matrices, i.e. the symmetric matrices of size $d$ with a positive spectrum.

We now give the proof of Proposition 1.
Proof. We write system (1.1) in matricial form:

$$
\partial_{t} Y+A(Y, \nabla) Y=0, \text { with } Y:=\left(\begin{array}{c}
\rho \\
v \\
\tilde{P}
\end{array}\right) \in \mathbb{R}^{1+d+d(d+1) / 2}
$$

and, if $\xi=\left(\xi_{i}\right)_{1 \leq i \leq d} \in \mathbb{R}^{d}$,

$$
A(Y, \xi):=\left(\begin{array}{ccc}
v \cdot \xi & \rho \xi^{T} & 0  \tag{2.1}\\
\frac{1}{\rho}\left(p^{\prime}(\rho) I_{d}+P\right) \xi & (v \cdot \xi) I_{d} & C(\xi) \\
0 & D(\xi) & (v \cdot \xi) I_{d(d+1) / 2}
\end{array}\right)
$$

When $d=2$, the symmetric matrix $P=\left(P_{i j}\right)_{1 \leq i, j \leq 2}$ can be identified as a vector $\tilde{P} \in \mathbb{R}^{3}$ given by

$$
\tilde{P}:=\left(\begin{array}{l}
P_{11} \\
P_{12} \\
P_{22}
\end{array}\right) .
$$

The matrices $C(\xi)$ and $D(\xi)$ are then given by

$$
C(\xi):=\left(\begin{array}{ccc}
\xi_{1} & \xi_{2} & 0 \\
0 & \xi_{1} & \xi_{2}
\end{array}\right) \text { and } D(\xi):=\left(\begin{array}{cc}
2 P_{11} \xi_{1}+2 P_{12} \xi_{2} & 0 \\
P_{21} \xi_{1}+P_{22} \xi_{2} & P_{11} \xi_{1}+P_{12} \xi_{2} \\
0 & 2 P_{12} \xi_{1}+2 P_{22} \xi_{2}
\end{array}\right)
$$

When $d=3$, the symmetric matrix $P$ can be identified as a vector $\tilde{P} \in \mathbb{R}^{6}$ :

$$
\tilde{P}:=\left(\begin{array}{l}
P_{11} \\
P_{12} \\
P_{13} \\
P_{22} \\
P_{23} \\
P_{33}
\end{array}\right) \text {. }
$$

The matrices $C(\xi)$ and $D(\xi)$ are then given by

$$
C(\xi):=\left(\begin{array}{cccccc}
\xi_{1} & \xi_{2} & \xi_{3} & 0 & 0 & 0 \\
0 & \xi_{1} & 0 & \xi_{2} & \xi_{3} & 0 \\
0 & 0 & \xi_{1} & 0 & \xi_{2} & \xi_{3}
\end{array}\right)
$$

and

$$
D(\xi):=\left(\begin{array}{ccc}
2(P \xi)_{1} & 0 & 0 \\
(P \xi)_{2} & (P \xi)_{1} & 0 \\
(P \xi)_{3} & 0 & (P \xi)_{1} \\
0 & 2(P \xi)_{2} & 0 \\
0 & (P \xi)_{3} & (P \xi)_{2} \\
0 & 0 & 2(P \xi)_{3}
\end{array}\right)
$$

where $(P \xi)_{i}$ denotes the $i$-th component of the vector $P \xi$.

We first show the implication $(2) \Rightarrow(1)$. Namely, if $p^{\prime}=0$ or $P=\lambda I_{d}$, then system (1.1) is Friedrichs-symmetrizable.

We thus suppose that $p^{\prime}=0$ or $P=\lambda I_{d}$.
Consider $S=S(Y) \in S_{n}^{++}(\mathbb{R})($ with $n=(d+1)(d+2) / 2)$ defined as a block matrix, compatible with $A$ :

$$
S=\left(\begin{array}{ccc}
1 & 0 & 0  \tag{2.2}\\
0 & S_{2} & 0 \\
0 & 0 & S_{3}
\end{array}\right), \text { with } S_{2} \in M_{d}(\mathbb{R}) \text { and } S_{3} \in M_{d(d+1) / 2}(\mathbb{R})
$$

Note that $S$ is symmetric definite positive if and only if the matrices $S_{2}$ and $S_{3}$ are also symmetric definite positive.

We can now compute the product $S A$ by block matrix mutiplication. We obtain that

$$
S A=\left(\begin{array}{ccc}
v \cdot \xi & \rho \xi^{T} & 0  \tag{2.3}\\
\frac{1}{\rho} S_{2}\left(p^{\prime}(\rho) I d+P\right) \xi & (v \cdot \xi) S_{2} & S_{2} C(\xi) \\
0 & S_{3} D(\xi) & (v \cdot \xi) S_{3}
\end{array}\right)
$$

Case $d=2$. Let us choose

$$
S_{2}:=\mu P^{-1} \text { and } S_{3}:=\mu\left(\begin{array}{ccc}
\frac{1}{2} q_{11}^{2} & q_{12} q_{11} & \frac{1}{2} q_{12}^{2} \\
q_{12} q_{11} & q_{11} q_{22}+q_{12}^{2} & q_{12} q_{22} \\
\frac{1}{2} q_{12}^{2} & q_{12} q_{22} & \frac{1}{2} q_{22}^{2}
\end{array}\right)
$$

where we denoted

$$
P^{-1}=\left(\begin{array}{ll}
q_{11} & q_{12}  \tag{2.4}\\
q_{12} & q_{22}
\end{array}\right) \in S_{2}^{++}(\mathbb{R})
$$

and

$$
\mu:= \begin{cases}\rho^{2} & \text { when } p^{\prime}(\rho)=0  \tag{2.5}\\ \rho^{2} \frac{\lambda}{p^{\prime}(\rho)+\lambda} & \text { when } P=\lambda I d\end{cases}
$$

The matrix $S_{2}$ is symmetric definite positive.
We see that the principal minors of $\mu^{-1} S_{3}$ are given by $M_{1}=\frac{1}{2} q_{11}^{2}>0, M_{2}=$ $\frac{1}{2} q_{11}^{2}\left(q_{11} q_{22}-q_{12}^{2}\right)>0$, and $M_{3}=\frac{1}{4}\left(q_{11} q_{22}-q_{12}^{2}\right)^{3}>0$ (recall that $P$ is positive definite). Hence by Sylvester's criterion, $S_{3}$ is definite positive, and $S$, defined by (2.2), is a positive definite matrix.

We compute that

$$
\begin{gather*}
\frac{1}{\rho} S_{2}\left(p^{\prime}(\rho) I d+P\right) \xi=\rho \xi  \tag{2.6}\\
S_{2} C(\xi)=\mu\left(\begin{array}{lll}
q_{11} \xi_{1} & q_{11} \xi_{2}+q_{12} \xi_{1} & q_{12} \xi_{2} \\
q_{12} \xi_{1} & q_{12} \xi_{2}+q_{22} \xi_{1} & q_{22} \xi_{2}
\end{array}\right),
\end{gather*}
$$

and, since $q_{i 1} P_{1 j}+q_{i 2} P_{2 j}=\delta_{i j}$ by 2.4,

$$
S_{3} D(\xi)=\mu\left(\begin{array}{cc}
q_{11} \xi_{1} & q_{12} \xi_{1} \\
q_{12} \xi_{1}+q_{11} \xi_{2} & q_{22} \xi_{1}+q_{12} \xi_{2} \\
q_{12} \xi_{2} & q_{22} \xi_{2}
\end{array}\right)=\left[S_{2} C(\xi)\right]^{T} .
$$

Hence for any $Y \in \mathcal{U}$, and for any $\xi \in \mathbb{R}^{2}$, the matrix $S(Y)$ is symmetric definite positive and (2.3) shows that the matrix $S(Y) A(Y, \xi)$ is symmetric. As a consequence, (1.1) is Friedrichs-symmetrizable when $d=2$.

Case $d=3$. Define again $S$ by equation 2.2 . Choose $S_{2}=\mu P^{-1}$ with $\mu$ as in (2.5), such that $S_{2}$ is symmetric definite positive and 2.6 holds again. Define $S_{3}$ by

$$
S_{3}:=\mu\left(\begin{array}{cccccc}
\frac{1}{2} q_{11}^{2} & q_{11} q_{12} & q_{11} q_{13} & \frac{1}{2} q_{12}^{2} & q_{12} q_{13} & \frac{1}{2} q_{13}^{2} \\
q_{11} q_{12} & q_{11} q_{22}+q_{12}^{2} & q_{11} q_{23}+q_{12} q_{13} & q_{12} q_{22} & q_{12} q_{23}+q_{22} q_{13} & q_{13} q_{23} \\
q_{11} q_{13} & q_{11} q_{23}+q_{13} q_{12} & q_{11} q_{33}+q_{13}^{2} & q_{12} q_{23} & q_{12} q_{33}+q_{23} q_{13} & q_{13} q_{33} \\
\frac{1}{2} q_{12}^{2} & q_{12} q_{22} & q_{12} q_{23} & \frac{1}{2} q_{22}^{2} & q_{22} q_{23} & \frac{1}{2} q_{23}^{2} \\
q_{12} q_{13} & q_{12} q_{23}+q_{13} q_{22} & q_{12} q_{33}+q_{13} q_{23} & q_{22} q_{23} & q_{22} q_{33}+q_{23}^{2} & q_{23} q_{33} \\
\frac{1}{2} q_{13}^{2} & q_{13} q_{23} & q_{13} q_{33} & \frac{1}{2} q_{23}^{2} & q_{23} q_{33} & \frac{1}{2} q_{33}^{2}
\end{array}\right),
$$

where we denoted again

$$
P^{-1}:=\left(\begin{array}{lll}
q_{11} & q_{12} & q_{13} \\
q_{12} & q_{22} & q_{23} \\
q_{13} & q_{23} & q_{33}
\end{array}\right)
$$

We see that $S_{3}$ is symmetric. Furthermore, the principal minors of $\mu^{-1} S_{3}$ are given by $M_{1}=q_{11}^{2} / 2>0, M_{2}=q_{11}^{2}\left(q_{11} q_{22}-q_{12}^{2}\right) / 2>0, M_{3}=q_{11}^{3} \operatorname{det}\left(P^{-1}\right) / 2>0$, $M_{4}=q_{11}\left(q_{11} q_{22}-q_{12}^{2}\right)^{2} \operatorname{det}\left(P^{-1}\right) / 4>0, M_{5}=\left(q_{11} q_{22}-q_{12}^{2}\right)^{2} \operatorname{det}\left(P^{-1}\right)^{2} / 4>0$ and $M_{6}=\operatorname{det}\left(P^{-1}\right)^{4} / 8>0$. Hence, by Sylvester's criterion, $S_{3}$ is definite positive and $S$ is symmetric definite positive.

We also check that

$$
S_{3} D(\xi)=\left[S_{2} C(\xi)\right]^{T}
$$

Hence equation (2.3) shows that $S A$ is symmetric, and, consequently, 1.1 is Friedrichs-symmetrizable when $d=3$.

We now show that $(1) \Rightarrow(2)$. Suppose that system 1.1$)$ is Friedrichs-symmetrizable, i.e. there is a matrix $S$ such that the product $S A$ is symmetric.

We write $S$ as a block matrix, compatible with $A$ :

$$
S:=\left(\begin{array}{ccc}
S_{1} & \alpha & \beta \\
\alpha^{T} & S_{2} & \gamma \\
\beta^{T} & \gamma^{T} & S_{3}
\end{array}\right)
$$

We compute the product $S A$ by block multiplication. Since $S A$ is symmetric, for any block of $S A$, denoted $(S A)_{i, j}(1 \leq i, j \leq 3)$, we must have $(S A)_{i, j}=(S A)_{j, i}^{T}$. For $i=j=3$, we obtain the constraint

$$
\gamma^{T} C(\xi)+(v \cdot \xi) S_{3}=\left[\gamma^{T} C(\xi)+(v \cdot \xi) S_{3}\right]^{T}
$$

Since $S_{3}$ is symmetric, we deduce that the product $\gamma^{T} C(\xi)$ is symmetric, for any $\xi \in \mathbb{R}^{d}$. By computing explicitely the product, we obtain that the only possibility is that $\gamma=0$.

For $i=1$ and $j=3$, we obtain the constraint

$$
[\alpha C(\xi)+(v \cdot \xi) \beta]^{T}=(v \cdot \xi) \beta^{T}+\frac{1}{\rho} \gamma^{T}\left(p^{\prime}(\rho) I d+P\right) \xi
$$

Since $\gamma=0$, we obtain that $\alpha C(\xi)=0$, for any $\xi \in \mathbb{R}^{d}$. By computing explicitely the product, we also obtain that $\alpha=0$. Hence $S$ has to be of the form

$$
S=\left(\begin{array}{ccc}
S_{1} & 0 & \beta \\
0 & S_{2} & 0 \\
\beta^{T} & 0 & S_{3}
\end{array}\right)
$$

For $i=2$ and $j=3$, we obtain the constraint

$$
\begin{equation*}
\rho \beta^{T} \xi^{T}+S_{3} D(\xi)=\left[S_{2} C(\xi)\right]^{T} \tag{2.7}
\end{equation*}
$$

Solving the linear system (2.7) for $S_{2}, S_{3}$ and $\beta$ gives after some computations that there exists two constants $\lambda_{1}, \lambda_{2}$ such that

$$
S_{2}=\lambda_{1} P^{-1} \text { and } \beta=\left\{\begin{array}{r}
\lambda_{2}\left(q_{11}, 2 q_{12}, q_{22}\right) \text { when } d=2  \tag{2.8}\\
\lambda_{2}\left(q_{11}, 2 q_{12}, 2 q_{13}, q_{22}, 2 q_{23}, q_{33}\right) \text { when } d=3
\end{array}\right.
$$

Note that it follows from these computations that $\beta D(\xi)=2 \lambda_{2} \xi^{T}$.
For $i=1$ and $j=2$ we obtain the constraint

$$
\begin{equation*}
\frac{1}{\rho} S_{2}\left(p^{\prime}(\rho) I d+P\right) \xi=\left[S_{1} \rho \xi^{T}+\beta D(\xi)\right]^{T}=\left(S_{1} \rho+2 \lambda_{2}\right) \xi . \tag{2.9}
\end{equation*}
$$

Equation (2.9) implies that $S_{2}$ is proportional to $\left(p^{\prime}(\rho) I d+P\right)^{-1}$. Since $S_{2}$ is also proportional to $P^{-1}$ by 2.8 , we obtain that the two matrices $P$ and $P+p^{\prime}(\rho) I d$ are proportional (recall that $S_{2}$ is invertible). Hence $p^{\prime}=0$ or $P=\lambda I d$.

## Declarations

- Conflict Of Interest statement: there is no competing interest.
- Data availability statement: there is no associated data.


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