

Detection of Seismic Events originating from Europa's Silicate Interior: Implications for constraining interior dynamics

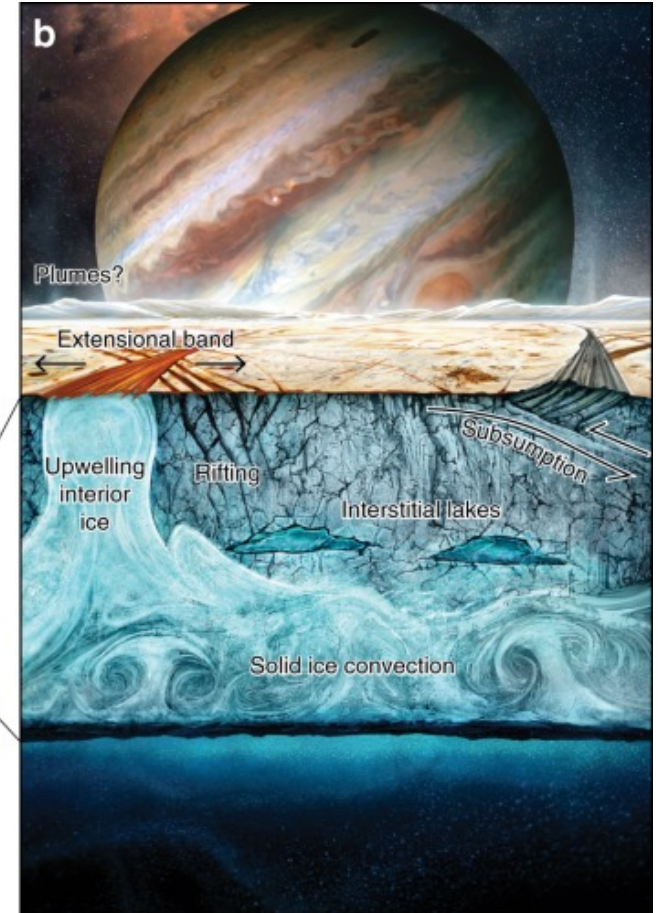
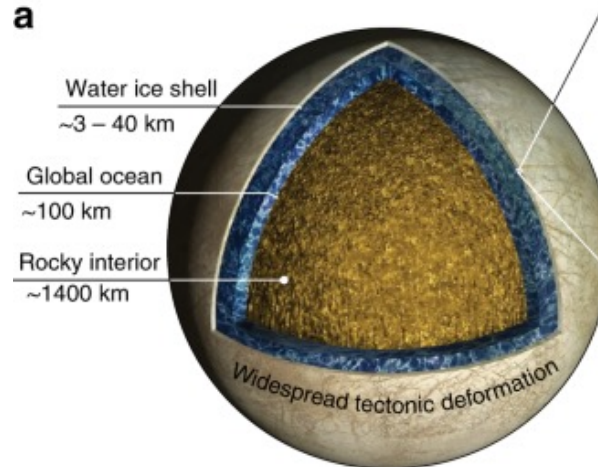
Angela G Marusiak
Postdoctoral Scholar



Jet Propulsion Laboratory
California Institute of Technology

Background: Europa Seismicity

- Signs of current geologic activity
- Tidal forces likely induce seismic activity
- Could point to ice shell dynamics



Implications for Deep Seismicity

- Hydrothermal systems?
- Volcanic activity?
- Geologic activity?
- Water-rock interactions
- Energy availability
- Habitable environments

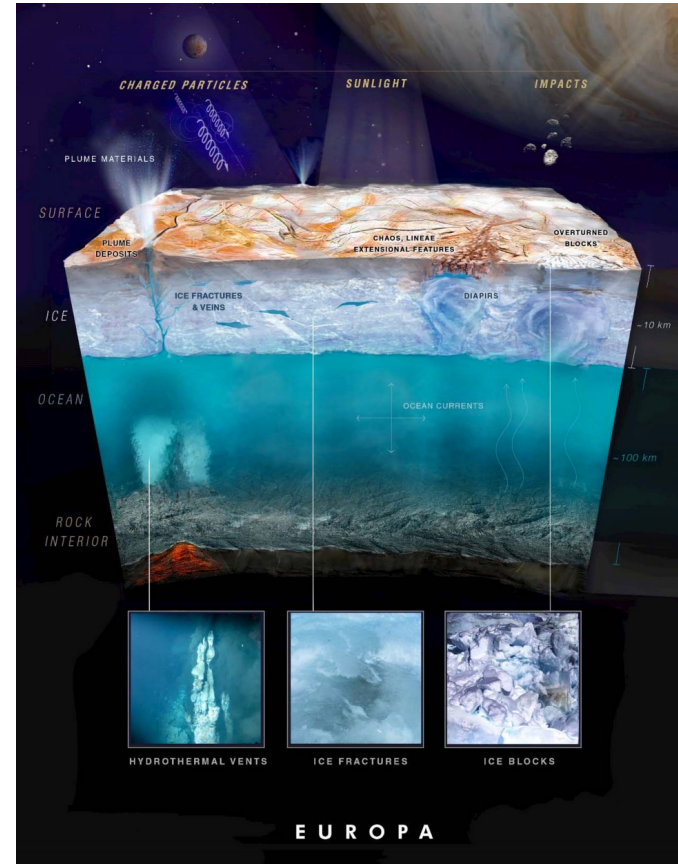
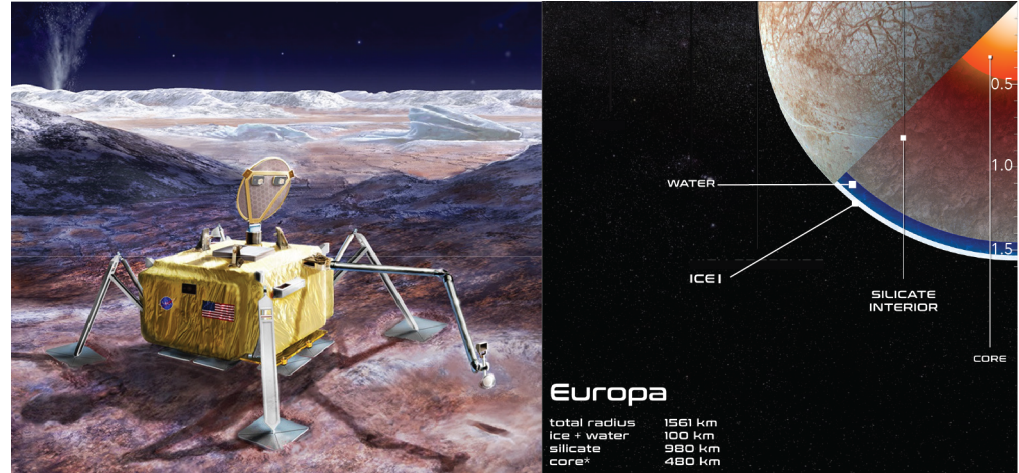


Image
credit:
NASA/JPL

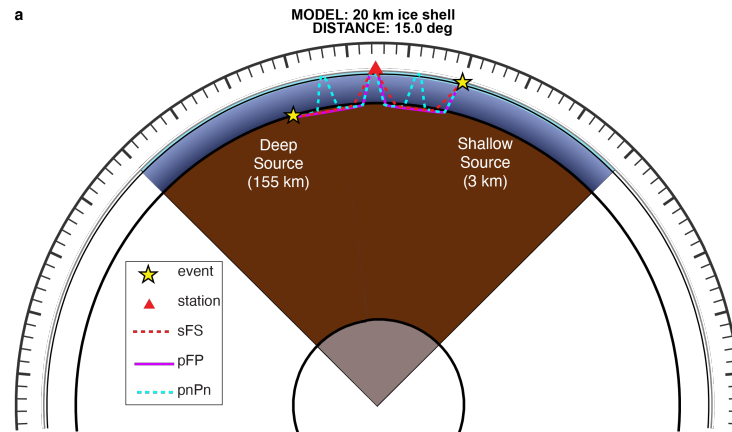
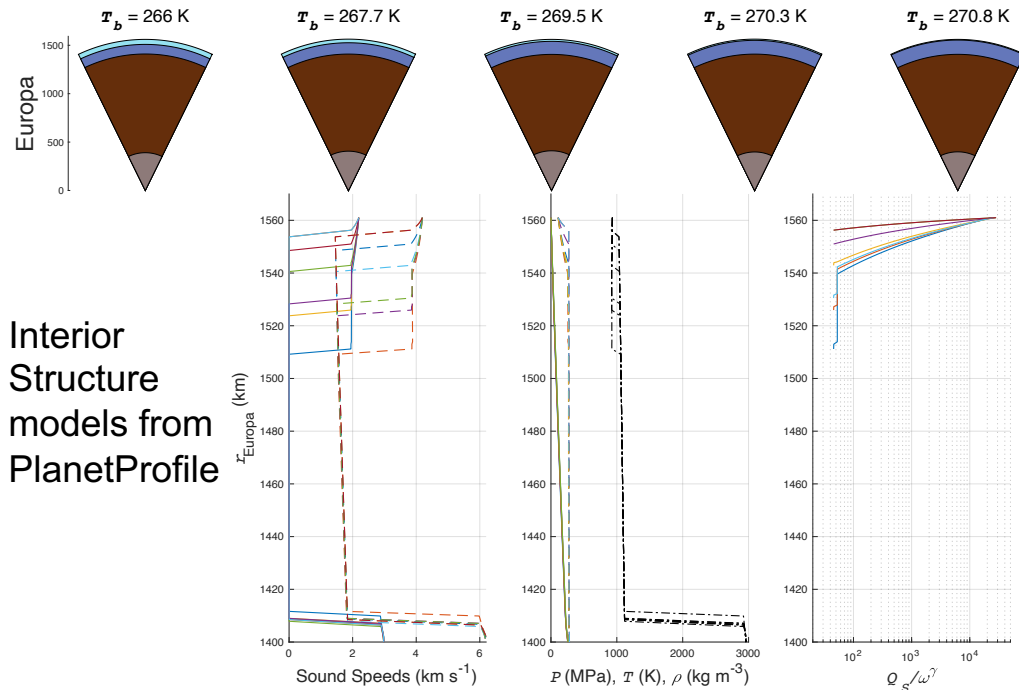
Background: Deep Interior

- Apollo seismometers recorded deep moonquakes from tidal forces
- Volcanic/hydrothermal activity has been theorized
- Geologic/tectonic state is unknown



Methods

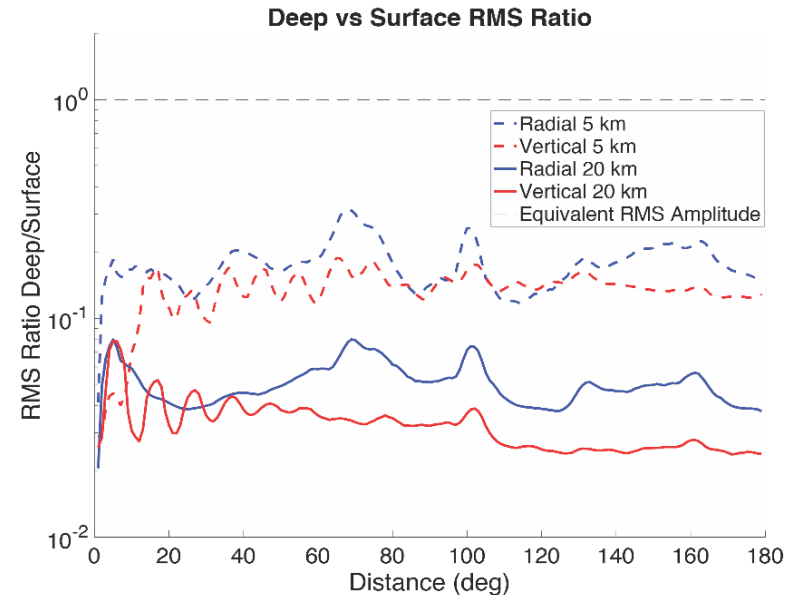
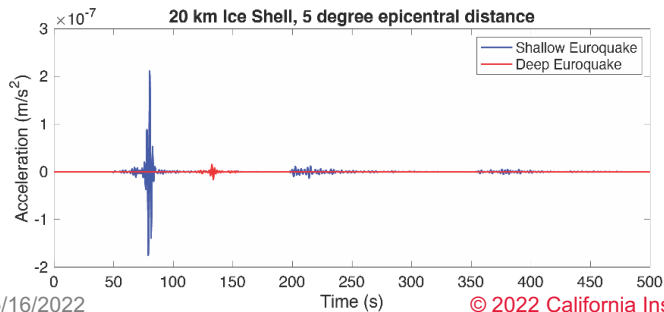
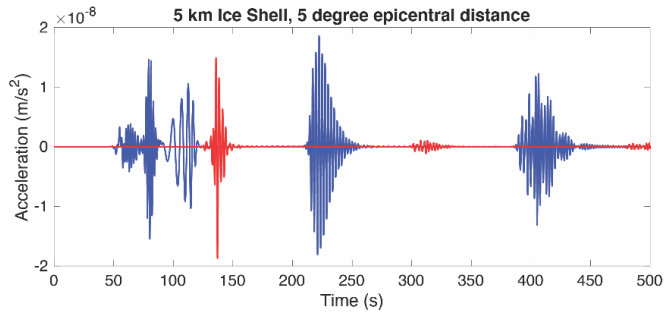
Let's see what what we can see



Waveforms from Instaseis and AxiSEM

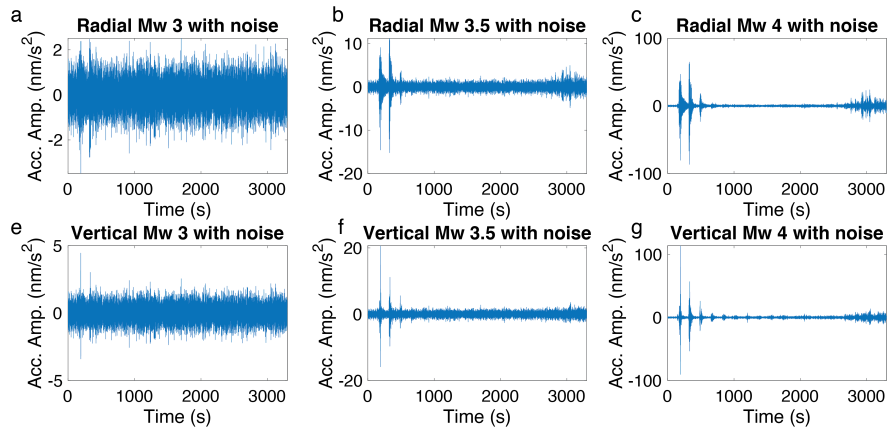
Surface vs Deep Waveforms

- Compare ground motion from 3 km deep event (shallow) and 155 km (deep)

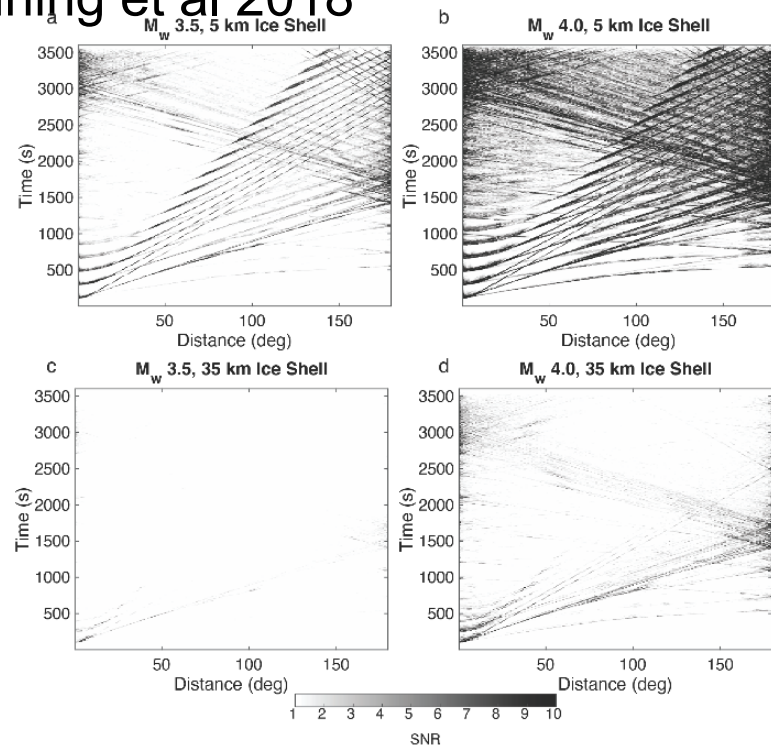


Signal-to-Noise Ratios

Background noise uses approach of Panning et al 2018



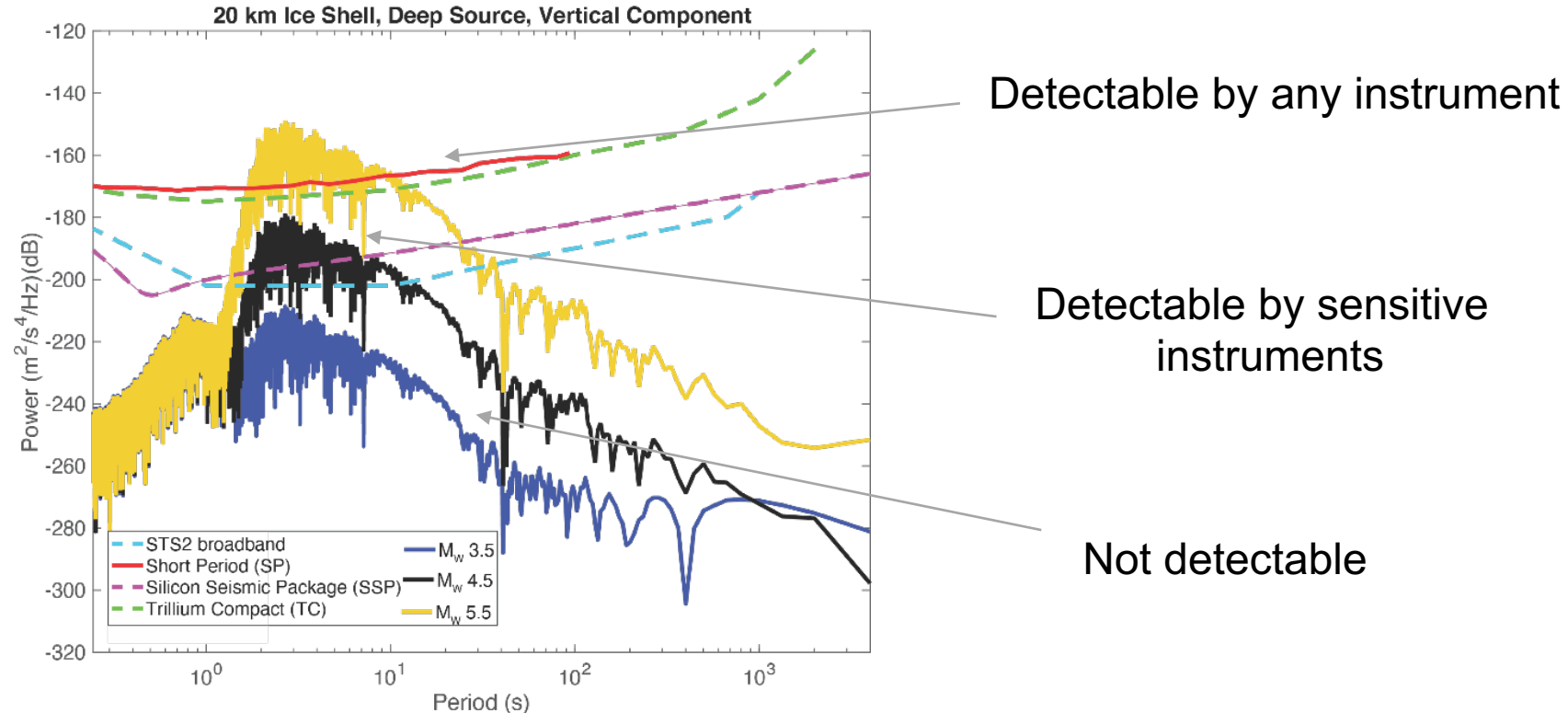
10 km ice shell



SNR depends on ice shell thickness

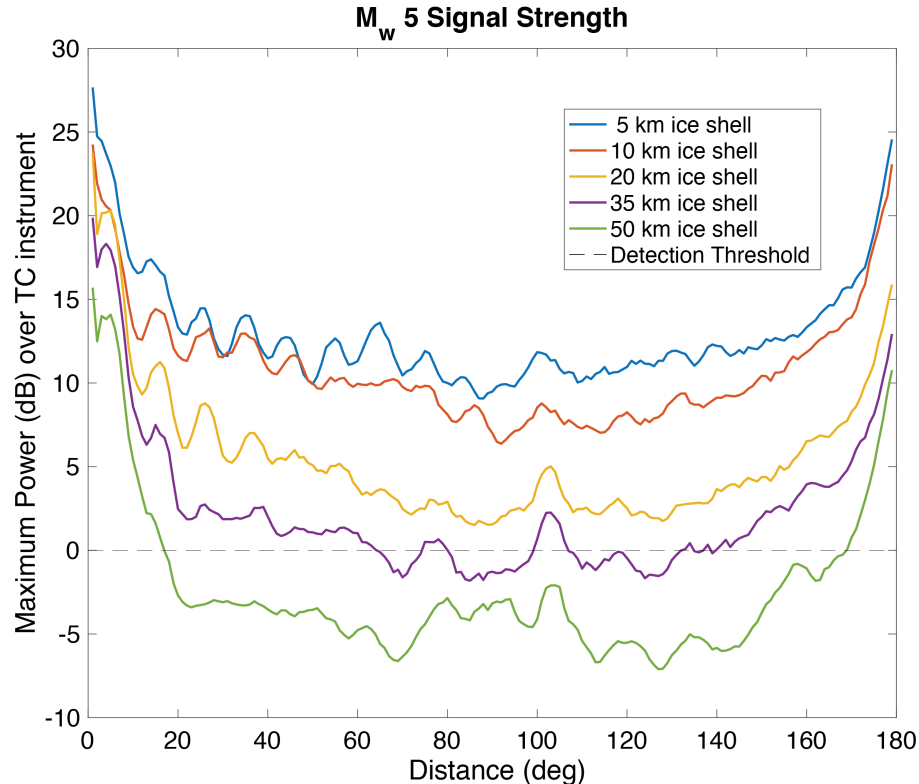
Instrument limits

What can we detect and see?



Ice Shell Thickness Effects

Measures maximum difference between event power and detection threshold



Mw 5.0 could be seen at any distance if ice shell is <35 km

Thicker ice shells need stronger ground motion or a nearby event

Summary

Link to
pre-print



SCAN ME

- Without instrument limits: Mw 3.5
- With sensitive instruments: Mw 4.0
- With less-sensitive instruments: Mw 5.0
- Thicker ice shells increase minimum magnitude
- **Lack of detection could be from instrument limits, not Europa**



Jet Propulsion Laboratory
California Institute of Technology

jpl.nasa.gov

Marusiak@jpl.nasa.gov