

Project Summary

Driving research question: What is the relationship between the parental melt compositions of the nakhlite and chassignite (martian) meteorites?

Why should we care? Elucidating the petrogenetic relationship of nakhrites and chassignites will constrain the degree of heterogeneity of the Martian mantle and crust.

What did we do? We conducted microanalysis on a suite of nakhrites and chassignites to determine their parental trapped liquid compositions.

What did we find? We found that nakhrites and chassignites likely formed from different, though chemically similar, parental melt compositions.

What is next? Trace element abundances will be determined for nakhlite and chassignite parental melts.

Methods

Electron Microprobe (EMP) Analysis:

- Inclusions: 15 kV accelerating voltage, 10 nA beam current
- Hosts: 15 kV accelerating voltage, 20 nA beam current
- Beam size: dependent on host phase, inclusion complexity [1]
- Inclusion phases are mode-normalized [2]

Melt Inclusion Rehomogenization:

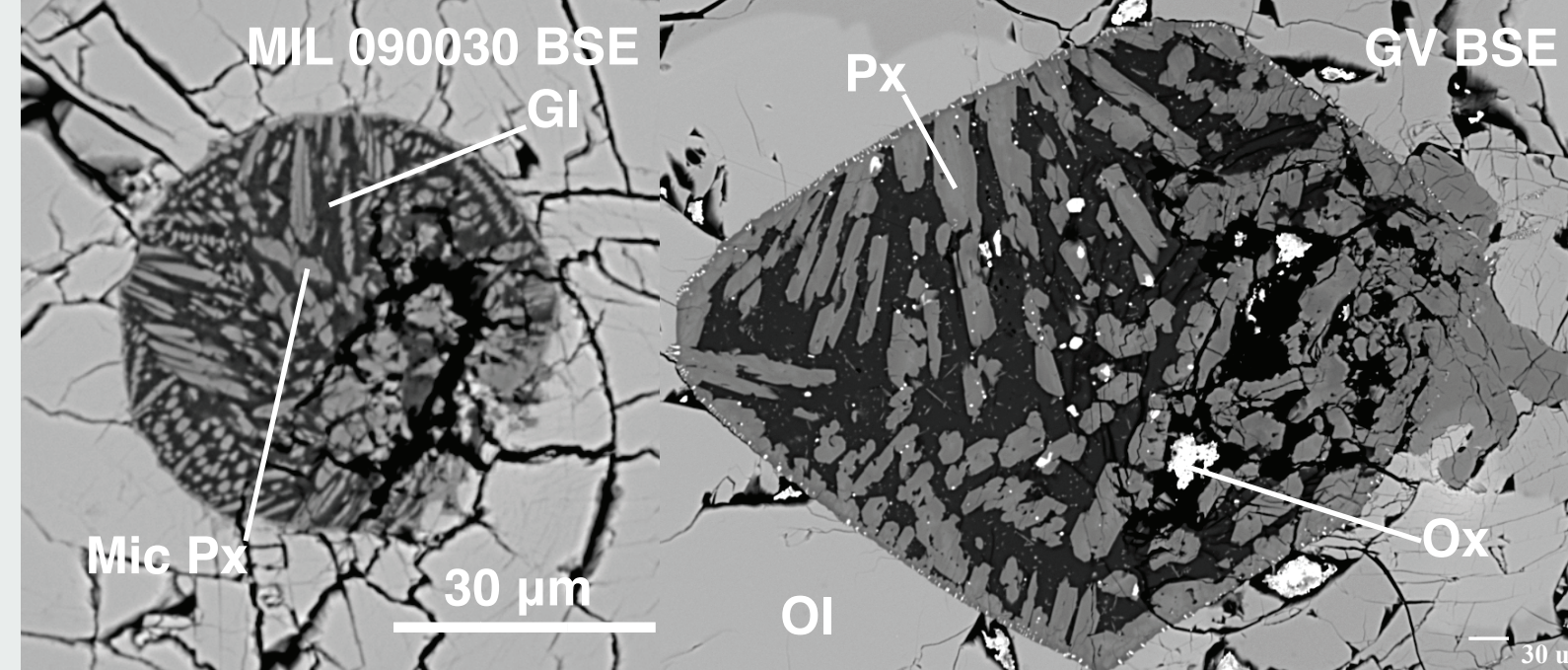
- Olivine-hosted: PETROLOG3 [3]
 - corrects reequilibration for Parental Trapped Liquid (PTL)
 - Required FeO_T value drawn from literature sources [4,5]
- Pyroxene-hosted: Rhyolite MELTS [6]
 - wall pyroxene added until it reaches equilibrium

Fractional Crystallization Modeling:

- Rhyolite MELTS [6]
 - FMQ buffer, 1 kbar pressure, H_2O from 0–2 wt. %

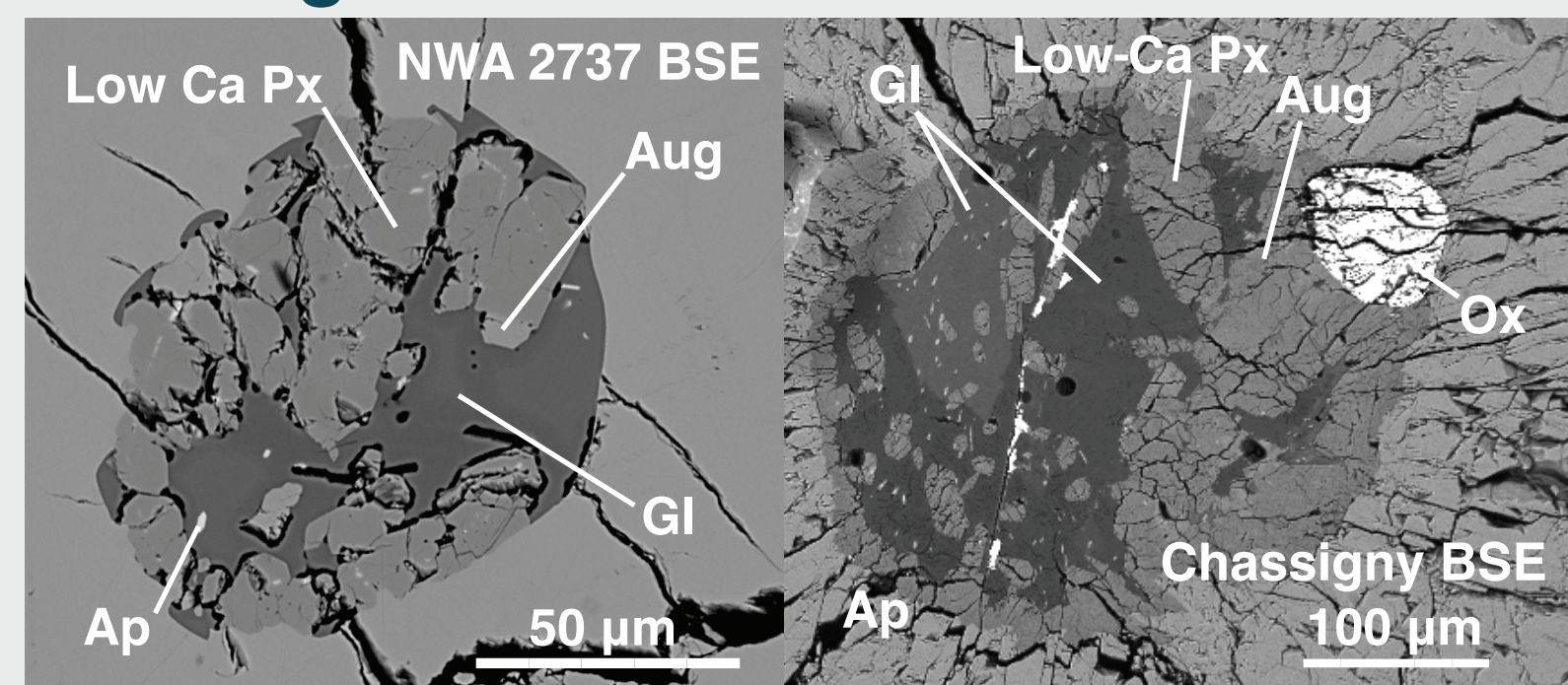
Melt Inclusions

Nakhlite Melt Inclusions:



Samples: (left) Miller Range 090030, simple phase assemblage, (right) Gobernador Valadares, complex phase assemblage.
Mic Px = Microlytic Pyroxene
Gl = Glass
Px = Pyroxene
Ol = Olivine
Ox = Oxide

Chassignite Melt Inclusions:

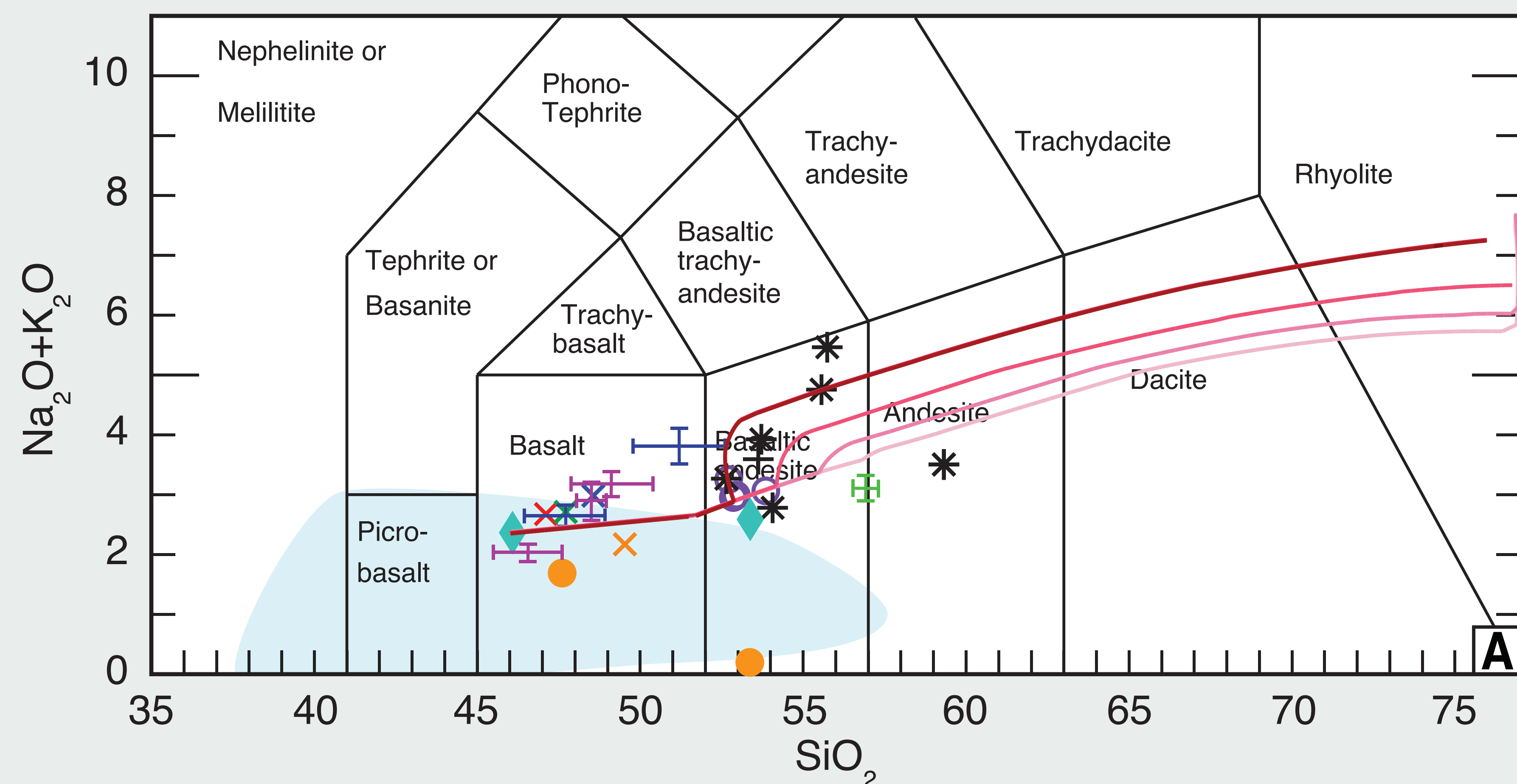


Samples: (left) Northwest Africa 2737 (right) Gobernador Valadares.
Low-Ca Px = Low Calcium Pyroxene
Aug = Augite
Ap = Apatite
Ox = Oxide

- Melt inclusions (MIs) are pockets of trapped melt inside of a host mineral [7]
- Host minerals are olivine and pyroxene (nakhrites) or olivine (chassignites)
- Phase assemblages range from simple (glassy) to complex (containing pyroxene, glass, oxides)

Results

Parental Trapped Liquid (PTL) Calculations and Fractional Crystallization Modeling:



Nakhlite PTLs

Pyroxene-Hosted MI

- + Miller Range 090030
- * Caleta el Cobre 022
- Northwest Africa 10645

Olivine-Hosted MI*

- + Miller Range 090030
- + Miller Range 090032
- + Gobernador Valadares

Chassignite PTLs

- Chassignite**
- ◆ Northwest Africa 2737***

Literature PTLs

- × Nakhla [4]
- × Miller Range 03346 [5]
- × Northwest Africa 2737 [10]
- × Nakhla [11]
- Nakhlite Bulk [8]

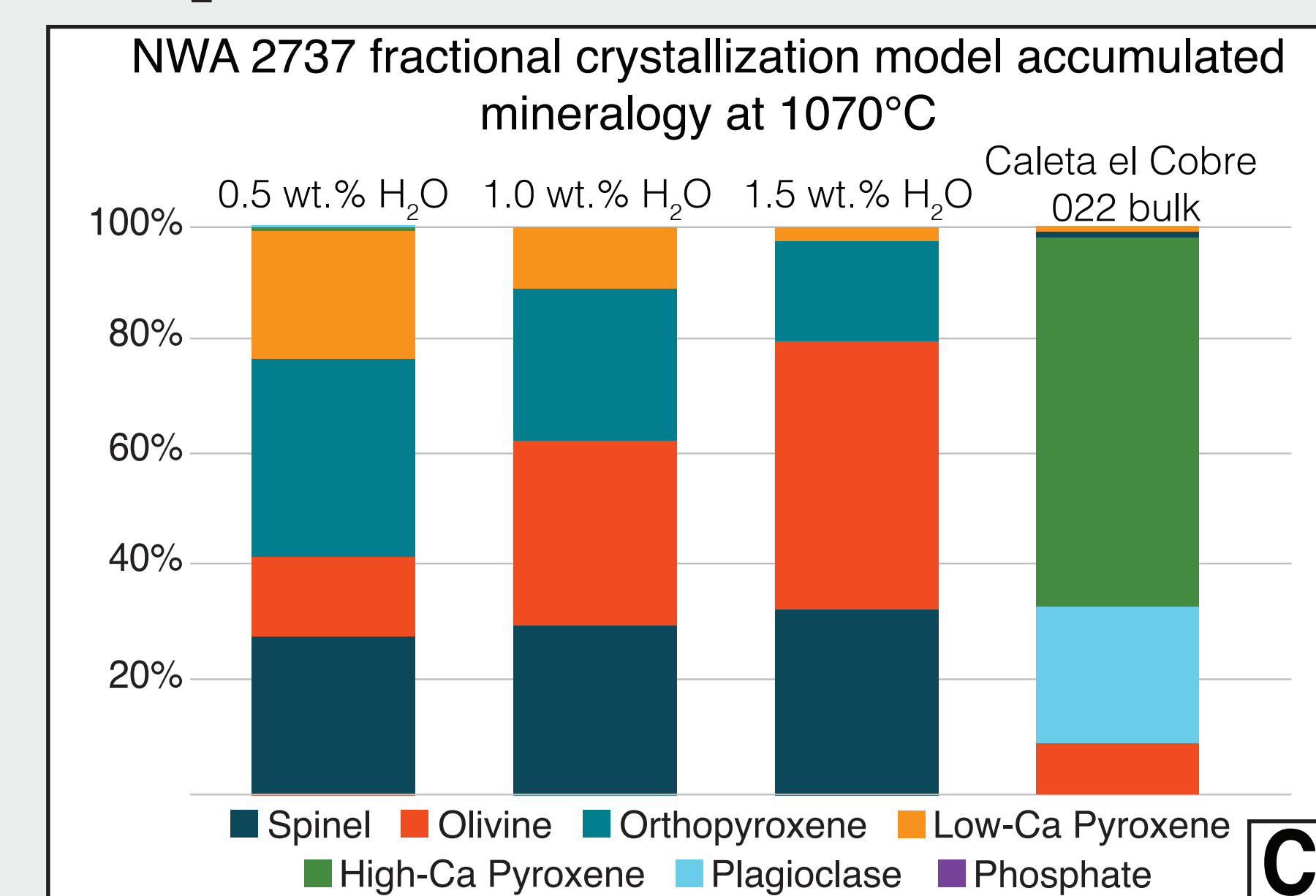
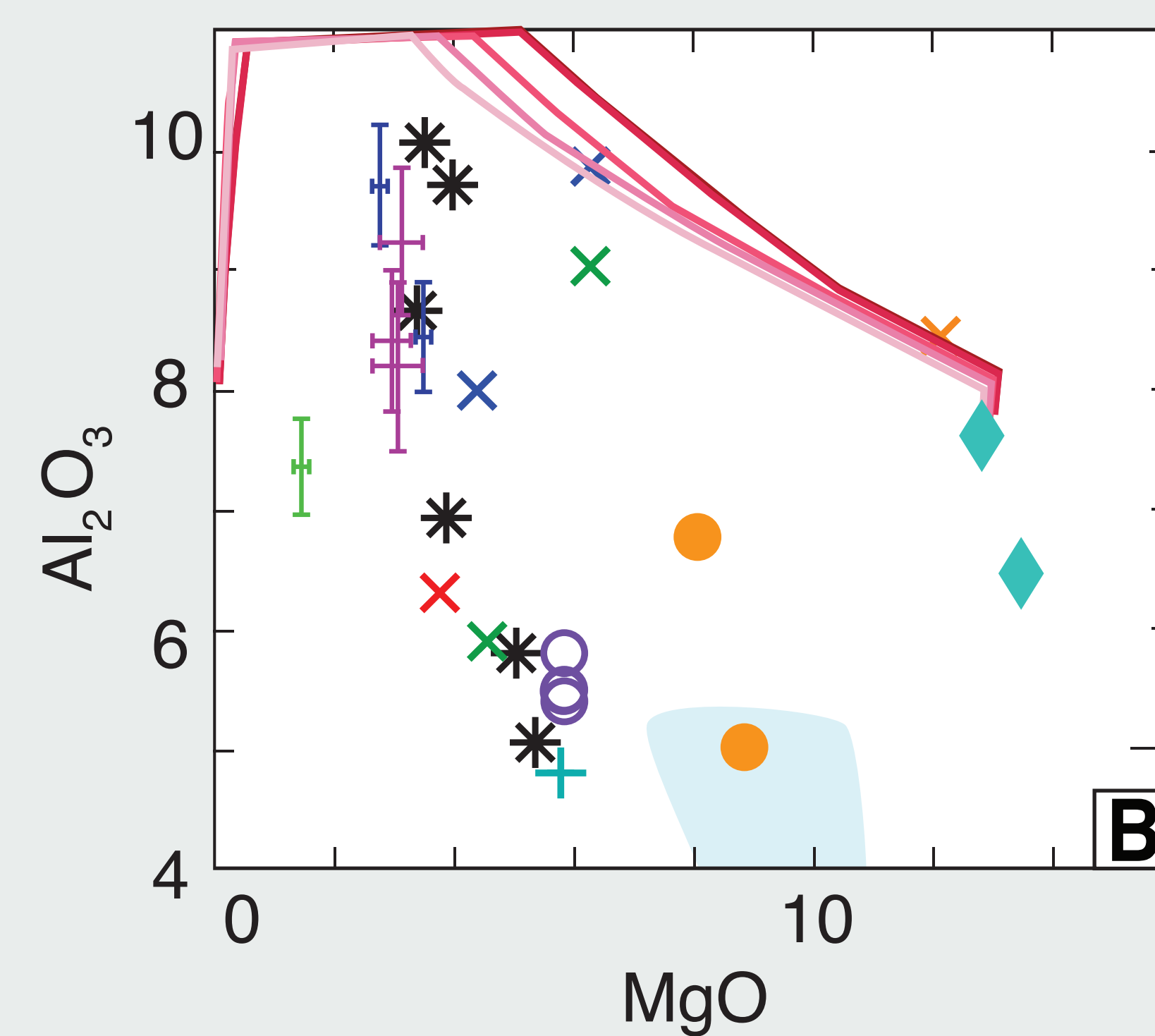
Model for NWA 2737 PTL

- 0.5 wt. % H_2O
- 1.0 wt. % H_2O
- 1.5 wt. % H_2O
- 2.0 wt. % H_2O

*Nakhlite olivine PTLs are calculated with FeO_T values from [4,5].

** The Chassignite PTL is calculated with an FeO_T from [12].

*** The Northwest Africa 2737 PTL is calculated with an FeO_T from [10].



- Nakhlite and chassignite parental trapped liquids (PTLs) share similar chemistry (Fig. A, B)
- Chassignite PTLs are more primitive than those in nakhrites (Fig. A)
- Northwest Africa (NWA) 2737 PTL liquid line of descent similar to pyroxene-hosted nakhlite PTLs in terms of K_2O and Na_2O , reaching 5.0 wt. % total with 49% liquid remaining at 1070°C (Fig. A)
- Overlap of nakhlite and chassignite PTLs in Al_2O_3 (2.7–10.3 wt. %) (Fig. B)
- NWA 2737 PTL crystallizes spinel, olivine, orthopyroxene + clinopyroxene, not observed mineralogy in nakhrites (clinopyroxene, olivine) (Fig. C)

Discussion

- There is a range in PTL compositions hosted in both nakhlite olivine and pyroxene
 - One phase may be xenocrystic [13]
- There is a large range of $\text{K}_2\text{O}+\text{Na}_2\text{O}$ values in both nakhlite and chassignite PTLs compared to bulk sample chemistry
 - Some samples may originate from a K_2O -enriched source
 - Not present in all samples: enrichment may be secondary
- Chassignite PTLs do not fractionate nakhlite mineralogy, but are similar in chemistry
 - Chassignites and nakhrites may arise from the same or similar sources
 - Chassignites and nakhrites do not share parental melt compositions

Future Work

- Analyze trace elements present in nakhlite and chassignite melt inclusions in order to characterize the source(s) of the suite
- Additional PTL calculations for nakhlite samples, such as Northwest Africa 998 and Northwest Africa 11013
- Constrain metasomatic or hydrothermal chemical influences on nakhlite and chassignite suite

References [1] Sonzogni, Y., and Treiman, A. (2015) Meteoritics & Planet. Sci., 50, 1880-1895. [2] Sarbadhikari, A.B., et al. (2017) Meteoritics & Planet. Sci., 52, 251-267. [3] Danyushevsky, L.V. and Plechov, P. (2011) Geochem Geophys Geosys, 12, No. 7. [4] Goodrich, C.A., et al. (2013) Meteoritics & Planet. Sci., 48, 2371-2405. [5] Imae, N., and Ikeda, Y. (2007) Meteoritics & Planet. Sci., 42, 171-184. [6] Gualda, G.A.R., et al., (2012) Journal of Petrology, 53, 875-890. [7] Schiano, P. (2003) Earth Science Rev., 63, 121-144. [8] Udry, A., and Day, J.M.D. (2018) GCA, 238, 292-315. [9] Sautter, V. et al., (2012) Meteoritics & Planet. Sci., 47, 330-344. [10] He, Q., et al. (2013) Meteoritics & Planet. Sci., 48, 474-492. [11] Stockstill, K.R., et al. (2005) Meteoritics & Planet. Sci., 40, 377-396. [12] McCubbin, F., et al. (2013) Meteoritics & Planet. Sci., 48, 819-853. [13] Ostwald, A.M., et. al. (2019) LPSC L, Abstract #1431.