

1 **The Role of Mesoscale Convective Systems in**
2 **Precipitation in the Tibetan Plateau region**

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7 **Key Points:**

- 8 • Mesoscale convective systems (MCSs) and associated precipitation are tracked in
9 the Tibetan Plateau (TP) region
10 • Long-lived and large MCSs produce substantial amounts of convective rainfall over
11 the Indo-Gangetic Plain
12 • Seasonal and extreme precipitation over the TP are dominated by smaller and short-
13 lived convective systems

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14 **Abstract**

15 Mesoscale convective systems (MCSs) have been identified as an important source
 16 of precipitation in the Tibetan Plateau (TP) region. However, the characteristics and
 17 structure of MCS-induced precipitation are not well understood. Infrared satellite im-
 18 agery has been used for MCS tracking, but cirrus clouds or cold surfaces can cause mis-
 19 classifications of MCS in mountain regions. We therefore combine brightness temper-
 20 atures from IR imagery with satellite precipitation data from GPM and track MCSs over
 21 the TP, at the boundary of the TP (TPB) and in the surrounding lower-elevation plains
 22 (LE) between 2000 and 2019. We show that MCSs are less frequent over the TP than
 23 earlier studies have suggested and most MCSs over land occur over the Indo-Gangetic
 24 Plain (LE) and the south of the Himalayas (TPB). In the LE and TPB, MCSs have pro-
 25 duced 10 % to 55 % of the total summer precipitation (10 % to 70% of summer extreme
 26 precipitation), whereas MCSs over the TP account for only 1 % to 10 % to the total sum-
 27 mer precipitation (1 % to 30 % of the total summer extreme precipitation). Our results
 28 also show that MCSs that produce the largest amounts of convective precipitation are
 29 characterized by longevity and large extents rather than by high intensities. These are
 30 mainly located south of the TP, whereas smaller-scale convection makes a greater con-
 31 tribution to total and total extreme precipitation over the TP. These results highlight
 32 the importance of convective scale modeling to improve our understanding of precipi-
 33 tation dynamics over the TP.

34 **Plain Language Summary**

35 Storm systems that extend over several hundred kilometers can produce substan-
 36 tial rainfall amounts over short periods. They are a risk for people’s life and livelihoods,
 37 as they may lead to flooding, extreme winds and heavy rainfall. The Tibetan Plateau
 38 (TP) has gotten increasing attention because it experiences drastic changes in the wa-
 39 ter cycle as a response to global warming. Although it is known that large storm sys-
 40 tems develop over the TP and in its populous surrounding regions, the rainfall charac-
 41 teristics from these storms are not well researched. It is not trivial to identify storm sys-
 42 tems in satellite images over high mountain regions, because high clouds and cold sur-
 43 face temperatures can give signals similar to those of storm clouds. We use therefore a
 44 new method to track large storm systems in satellite images over the TP, in order to clar-
 45 ify their role in the water cycle and extreme rainfall. Our results show that the storms
 46 that produce most heavy rainfall occur over the Indo-Gangetic Plain and south of the
 47 Himalayas. Storm systems over the TP are generally much smaller in size and shorter
 48 in duration, which means that climate model simulations at high spatial resolutions are
 49 needed to further investigate these.

50 **1 Introduction**

51 Mesoscale convective systems (MCSs) are organized convective storm complexes,
 52 which extend over several hundreds of kilometers and produce large areas of convective
 53 and stratiform precipitation (Houze, 2004). MCSs have more complex dynamics than
 54 unicellular convective storms, but are primarily defined by their spatial extent (Houze,
 55 2004). Many different forcings can drive mesoscale organization of convection. Thus, the
 56 structure and precipitation characteristics of an MCS can take different forms depend-
 57 ing on the region of genesis. Weather systems that are usually classified as MCSs include
 58 tropical cyclones, squall lines, lake-effect snow events and polar lows. In the continen-
 59 tal midlatitudes, MCS formation is often related to mountain flow dynamics and MCSs
 60 are often found in or close to mountain regions. On the leeward side of the Rocky Mountains
 61 (over the Great Plains) (Hitchcock et al., 2019; Cheeks et al., 2020; H. Hu et al., 2020),
 62 in the West-African Sahel (Redelsperger et al., 2002; Vondou et al., 2010; Klein et al.,

2018) and in the European Alps (Morel & Senesi, 2002; Feidas, 2017), MCSs produce a significant portion of the total precipitation in a season and can lead to severe weather.

The Tibetan Plateau (TP) covers an area of two and a half million square kilometers and is the world’s most extensive mountain region. The headwaters of most of Asia’s major rivers rise in the mountains of the TP, and their discharge regimes are mainly dominated by monsoonal precipitation (Zhang et al., 2013). An outstanding characteristic of the TP is its high elevation and steep topography, which results in complex interactions between local mountain features and large-scale atmospheric dynamics. Studies have identified MCSs to be one of the most important precipitation-producing mechanisms over the TP (Tao & Ding, 1981; B. Wang, 1987; Li et al., 2008; Sugimoto & Ueno, 2012). Some extreme rainfall and flood events in the populated downstream regions to the south and east of the TP have been attributed to MCSs and to mesoscale vortices that formed over TP (Yasunari & Miwa, 2006; Shi et al., 2008; Xiang et al., 2013; Rasmussen & Houze Jr, 2012). This demonstrates that MCSs can pose a direct threat to life, to people’s livelihoods, to crop yields and to infrastructure. In addition, MCSs play an important role in the hydrological cycle, as they may account for a significant amount of the annual or seasonal rainfall, for example in North America (Fritsch et al., 1986). Convection-permitting model simulations project increases in MCS intensity for some regions of the globe (Prein, Liu, Ikeda, Trier, et al., 2017; Fitzpatrick et al., 2020) and convective precipitation is likely to increase at much larger rates than the average precipitation increase by 7 % per degree of warming according to the Clausius-Clapeyron relation (Berg et al., 2013; Ban et al., 2015). It is therefore crucial that we understand the scales and formation processes for extreme convective precipitation, particularly for mountain regions like the TP, which are likely to experience drastic environmental changes due to global warming (Bibi et al., 2018; Yao et al., 2019).

Despite the fact that much literature on convection in the TP region focuses on MCSs, the main drivers behind the systems and their significance for current and future precipitation regimes are not well understood. It is not clear how characteristics of MCSs that initiate over the TP differ from those of the monsoon-related convective systems that occur south of the Himalayas (Houze et al., 2007; Romatschke et al., 2010; G. Chen & Kirtman, 2018). A number of studies that researched MCS at elevations higher than 3000 m above sea level (a.s.l.) have identified the central and eastern parts of the TP as the main source regions for MCS genesis (Jiang & Fan, 2002; Li et al., 2008; Sugimoto & Ueno, 2012; L. Hu et al., 2017). However, due to the difficulty in attributing precipitation events to specific storm systems, the role of MCSs in precipitation could only roughly be estimated. In addition, convective cells can be misclassified when infrared (IR) satellite imagery is used to track MCSs in a high mountain region and image scenes include cold cirrus cloud tops (Rossow & Schiffer, 1999; L. Hu et al., 2017) or cold surfaces under clear-sky conditions that result in a similar IR brightness temperature (Esmaili et al., 2016). Observation-based studies of MCSs are therefore more limited than model studies for the TP region and most cover only a few years, because high-resolution satellite records of more than decade’s length have only recently become available. It is crucial that we establish an accurate climatology of MCSs and understand their role in precipitation so that we can advance our knowledge about precipitation dynamics in the TP region. The effect of MCSs on precipitation is key to an improved understanding of the drivers and scales for extreme precipitation which in turn is necessary for more accurate estimations of future changes to precipitation regimes and extreme events.

The aim of this study is to describe the characteristics of MCSs in the TP region using a novel tracking method to interpret satellite observations covering the past two decades (from 2000 to 2019). To provide a broad overview of different types of MCSs, we compare MCSs over the TP with MCSs that cross the TP boundary (TPB) and MCSs that develop over the surrounding lower-elevation (LE) regions. We focus on the struc-

115 ture and characteristics of MCS-induced precipitation and the contribution of MCSs to
 116 seasonal and extreme rainfall.

117 The manuscript is organized as follows. In section 2, we describe the tracking al-
 118 gorithm and datasets used in this study and also summarize results from a pilot study
 119 in which we tested the sensitivity of the tracking to different thresholds and criteria. In
 120 section 3, we present the results including the spatial and temporal characteristics of MCS
 121 tracks and the precipitation features associated with them. The role of MCSs for pre-
 122 cipitation in the TP region and some possible mechanisms for MCS formation are dis-
 123 cussed in section 4. Finally, a summary and the main conclusions are given in section
 124 5.

125 2 Data and Methods

126 2.1 Tracking algorithm

127 One of the most commonly used methods to identify MCSs is to detect a contigu-
 128 ous area of cloud top temperature minima in IR satellite imagery. A specific type of MCS
 129 is a so-called *Mesoscale convective complex* (MCC), originally defined by Maddox (1980).
 130 MCC is a cloud system with a continuous area of at least 100 000 km², within which the
 131 maximum temperature is -32°C (241 K) and which includes a region of at least 50 000
 132 km², within which the maximum temperature is -52°C (221 K). An additional criterion
 133 is that these two conditions must persist for at least six hours for a MCC to be identi-
 134 fied. A myriad of studies has followed an approach similar to this for global and regional
 135 MCS tracking, mostly using cloud top temperature thresholds of between 230 K and 255
 136 K over minimum areas that range from 2000 km² to 50 000 km² (Sugimoto & Ueno, 2012;
 137 Esmaili et al., 2016; G. Chen & Kirtman, 2018; Klein et al., 2018; Cheeks et al., 2020).
 138 Most observational MCS studies in the TP region have used global databases for track-
 139 ing (Li et al., 2008) or thresholds that are also used in global analyses for MCS identi-
 140 fication (Guo et al., 2006). However, using universal thresholds can be problematic in
 141 a mountain environment like the TP and can lead to low surface temperatures from high
 142 mountain tops being confused with high cloud tops from deep convective clusters, par-
 143 ticularly at night and during the winter. This has, for example, been discussed in Esmaili
 144 et al. (2016), who presented a global cloud cluster tracking with unrealistically high amounts
 145 of cloud clusters over the TP during winter. The atmospheric transmittance at wave-
 146 lengths corresponding to the IR channels used for tracking ($\sim 10.8 \mu\text{m}$) is relatively high,
 147 while surface emissivity at these wavelengths is generally low for dry regions (Schädlich
 148 et al., 2001). This means that retrieved clear-sky brightness temperatures are lower than
 149 the actual surface temperatures and that there is an additional risk to confuse cold sur-
 150 faces with high cloud tops. Another risk of exclusively using cloud top temperatures as
 151 a proxy for convective activity in mountain regions is that convective systems can also
 152 be confused with cirrus cloud shields that are not necessarily the remnants of a storm
 153 system. Kukulies et al. (2019) showed that cirrus clouds are among the most frequent
 154 cloud types over the central and southern parts of the TP between May and September.
 155 To address the above named issues, we have thus used a new objective tracking method
 156 that reduces misclassifications of MCS attributable to cirrus cloud layers and cold sur-
 157 faces.

158 We use half-hourly satellite precipitation estimates from the Global Precipitation
 159 Measurement Mission (GPM) in combination with brightness temperatures from IR im-
 160 agery. The brightness temperatures used in this study are from merged, angle-corrected
 161 IR observations acquired by sensors on board Meteosat, GMS/Himawari, Meteosat and
 162 GOES. The NCEP/CPC (National Centers for Environmental Prediction/Climate Pre-
 163 diction Center) dataset can be downloaded from NASA GES DISC ([https://disc.gsfc](https://disc.gsfc.nasa.gov/datasets/GPM_MERGIR_V1/summary)
 164 [.nasa.gov/datasets/GPM_MERGIR_V1/summary](https://disc.gsfc.nasa.gov/datasets/GPM_MERGIR_V1/summary)) at 30 minute resolution and with 4km
 165 grid spacing. In order to obtain the same spatial resolution as the satellite precipitation

Table 1. Criteria for different tracking methods

| Test | Threshold | Minimum Extent | Time Requirement |
|-----------------------|--|--|--|
| T_b | ≤ 230 K | 100 grid cells | ≥ 6 hrs |
| T_b cold core | ≤ 230 K < 200 K | 100 grid cells 1 grid cell | ≥ 6 hrs 1/lifetime |
| T_b heavy rain core | ≤ 230 K < 200 K ≥ 5 mm h ⁻¹ | 100 grid cells 1 grid cell 10 grid cells | ≥ 6 hrs 1/lifetime 1/lifetime |
| <i>Precip</i> | ≥ 5 mm h ⁻¹ | 10 grid cells | ≥ 6 hrs |
| <i>TCS</i> | ≤ 230 K < 200 K ≥ 5 mm h ⁻¹ | 20 grid cells 1 grid cell 5 grid cells | ≥ 3 hrs 1/lifetime 1/lifetime |

*The detection and tracking criteria are modified from Yuan and Houze Jr (2010) and G. Chen and Kirtman (2018)

166 data product GPM IMERG V06 (https://disc.gsfc.nasa.gov/datasets/GPM_3IMERGHH
167 [_06/summary?keywords=GPM%20IMERG%20v06%2030%20MIN](https://disc.gsfc.nasa.gov/datasets/GPM_3IMERGHH_V06_summary?keywords=GPM%20IMERG%20v06%2030%20MIN)), which has a spatial res-
168 olution of 0.1 °, we regridded the brightness temperature data to match this. The track-
169 ing was performed in 30 minute time steps to match the original temporal resolution of
170 both datasets, for the period 2000 to 2019.

171 The main difference between our tracking method and more conventional track-
172 ing methods is that we use lower temperature thresholds and require that cloud cells de-
173 velop a cold core that drops below 200K at least once during their lifetime and that at
174 least 10 % of the minimum cloud area (10 pixels) correspond to a rain rate of at least
175 5 mm h⁻¹. The purpose of these extra criteria is to avoid high cirrus clouds or strati-
176 form cloud layers with lower rain rates being mistakenly identified as MCSs. By com-
177 bining IR and precipitation, we increase the captured lifetime of the evolving cloud cells
178 (more details are given in the following section).

179 Using the python package *tobac* (Heikenfeld et al., 2019), cloud features are iden-
180 tified at each time step using Gaussian filtered brightness temperatures (T_b) and then
181 linked over time based on their location and propagation speed. More details about this
182 linking method can be found in Heikenfeld et al. (2019). We define a cloud feature as
183 a contiguous area of at least 100 pixels (about 10 000 km²) in which brightness temper-
184 atures do not exceed 230 K (Table 1). With this minimum area threshold we focus on
185 MCS at the upper bounds of the *mesoscale* (Orlanski, 1975), because we assume that
186 these are the MCS with the largest impacts and the MCS which are more likely to in-
187 teract with the large-scale atmospheric circulation. Additionally, previous studies have
188 identified mesoscale vortices with horizontal dimensions between 100 - 1000 km over the
189 TP (B. Wang, 1987; Tao & Ding, 1981; Yasunari & Miwa, 2006), so this study allows
190 us to check whether cloud cells over the TP can develop at these scales. Each detected
191 cloud feature is then augmented by including adjacent pixels with brightness temper-
192 atures up to 245 K (235 K for December, January and February). This allows us to in-

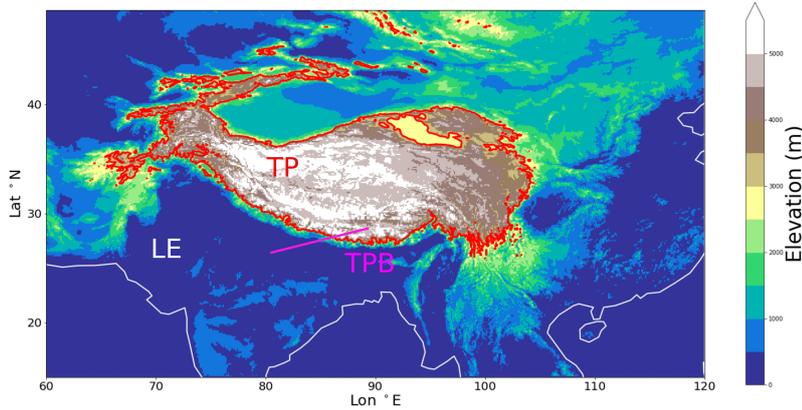


Figure 1. Study area (15 – 50 °N, 60 – 120 °E) for regional MCS tracking. The colors show elevations [m a.s.l.] and the red line indicates the 3000 m boundary of the TP. MCS are divided into systems located outside of this boundary (LE), systems over the TP and systems, which cross the 3000 m boundary during their lifetime (TPB).

193 clude as much of the cloud cell as possible, while excluding surrounding cold surfaces.
 194 The thresholds for selecting additional pixels to augment detected cloud features were
 195 chosen by evaluating distributions of hourly surface temperatures from ERA5 ([https://](https://cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5-single-levels?tab=overview)
 196 [cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5-single-levels](https://cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5-single-levels?tab=overview)
 197 [?tab=overview](https://cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5-single-levels?tab=overview)) and by performing test runs with different thresholds. Figure 2 shows
 198 that our cloud top temperature threshold of 245 K (235 K during winter), it can be does
 199 not overlap with the corresponding seasonal distribution of hourly surface temperatures.
 200 Once cloud features have been detected and augmented for each time step, they are linked
 201 over time and a cloud system must persist for at least 6 hours to be retained as a po-
 202 tential MCS. Due to computational resources linking of features was performed on yearly
 203 aggregated files, which means that MCS at the boundary between two years appear as
 204 separate tracks. Finally, two additional criteria are applied. To be classified as an MCS,
 205 the feature must have an area of at least one pixels where the maximum temperature
 206 does not exceed 200 K and an area of at least 10 pixels ($\sim 10\%$ of minimum cloud area)
 207 with rain rates of at least 5 mm h^{-1} at least once during their lifetime (T_b heavy rain
 208 core in Table 1).

209 It should be noted that the study area (15 – 50 ° N, 60 – 120 °E) encompasses re-
 210 gions with substantially different precipitation regimes, such as the Indo-Gangetic Plain,
 211 which are dominated by monsoon depressions (Hurley & Boos, 2015) and the generally
 212 drier TP (Fig. 1). Considering such a wide area with diversified background climates pro-
 213 vides a regional overview of MCSs, allowing those over the TP to be compared with those
 214 initiated over more populous regions. In this study, we distinguish between three classes
 215 of systems: MCSs and precipitation events that remain within the 3000 m boundary of
 216 the plateau (TP), MCS that cross the 3000 m boundary with at least 10 pixels once dur-
 217 ing their lifetime (TPB) and MCSs and precipitation events outside the 3000 m bound-
 218 ary at lower elevations (LE). These three subregions are visualized in Figure 1.

219 Figure 3 shows brightness temperatures and precipitation data for the study area.
 220 The snapshot shows a mature MCS on July 20th, 2008 and the succeeding plots show
 221 the evolution of the MCS track. The black line indicates the center of the MCS for ear-

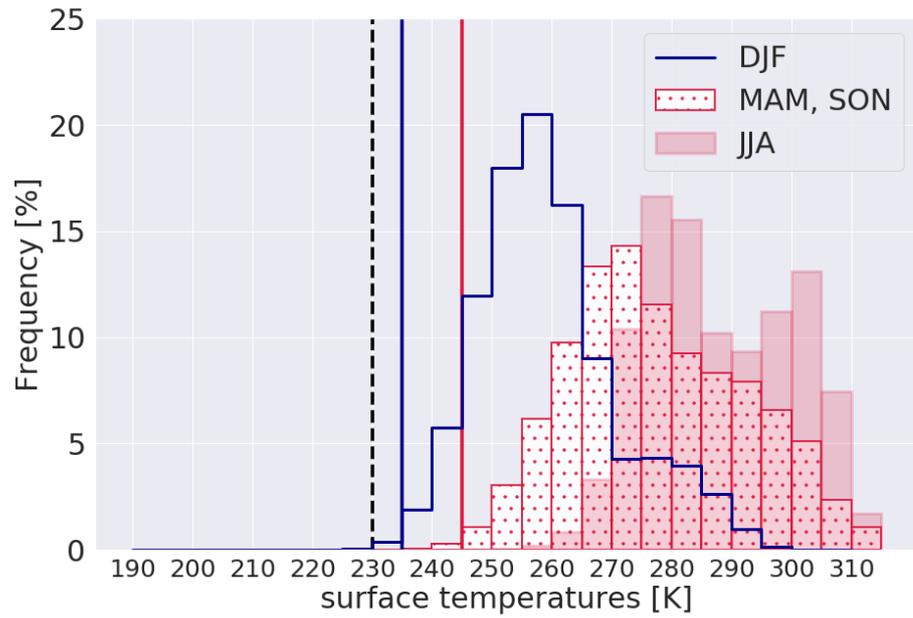


Figure 2. Seasonal distributions of hourly surface temperature over the TP (for surface elevations ≥ 3000 m) based on ERA5 data for 2000 to 2019. The dashed black line shows the threshold used for detection of cloud features and the vertical red and blue lines indicates show the thresholds used to select pixels for augmentation of cloud features in winter (December to February) and outside of the winter season (March to November), respectively.

222 lier time steps and the red dot indicates the center of the MCS at the time of the snap-
 223 shot. This tracked MCS produced substantial amounts of heavy rainfall in the downstream
 224 region to the east of the TP. We use this well-known event, which also coincided with
 225 a mesoscale disturbance in vorticity (Curio et al., 2019), as a case study to prove that
 226 our tracking algorithm is able to capture the known evolution of the system.

227 2.2 Comparison of tracking methods

228 It should be noted that the atmospheric variable selected as a proxy for convec-
 229 tion (e.g. cloud top temperature, precipitation, vorticity, wind speed) determines the spa-
 230 tial and temporal characteristics of the tracked MCS. There are many advantages to us-
 231 ing precipitation, as it is a key component in the water cycle that has direct impacts on
 232 hydrology and society. It is also reasonable straightforward to compare precipitation tracks
 233 with model and reanalysis data, whereas cloud top temperatures are usually not avail-
 234 able as standard products. However, the part of an MCS in which precipitation is pro-
 235 duced is usually smaller and more short-lived than the cloud system as a whole. Hence,
 236 using precipitation as a proxy for convection provides a more limited view of both the
 237 structure and evolution of tracked storm systems compared to cloud top temperature.
 238 This motivated us to perform a pilot study, in which we compare four algorithms with
 239 different tracking criteria to improve our understanding of the implications of different
 240 tracking methods for MCSs in the TP region.

241 Table 1 summarizes the criteria for the four different algorithms. First, we performed
 242 tracking using only brightness temperatures (T_b) applying the same criteria to these as
 243 we used in our combined tracking for cloud feature identification. We then added the
 244 cold core criterion (in T_b cold core) and the heavy rain criterion (in T_b heavy rain core)
 245 from our combined tracking (described in section 2.1). Finally, we implemented track-
 246 ing based on areas of contiguous precipitation using different rain rate thresholds. The
 247 results for the run with a threshold of 5 mm h^{-1} are presented here, because it repre-
 248 sents the best compromise for which precipitation cells could be tracked in the more hu-
 249 mid parts of the study domain as well as over the drier TP. In addition, this threshold
 250 lies within the range of rainfall intensities that are typically used to classify convective
 251 precipitation (Gaál et al., 2014) and has also been used by other MCS studies (Prein,
 252 Liu, Ikeda, Bullock, et al., 2017).

253 For comparison, we also implemented tracking for convective systems within the
 254 TP (elevation is equal or above 3000 a.s.l.), using lower minimum thresholds for area (\geq
 255 20 pixels) and persistence (\geq 3 hours) (TCS in Table 1). These smaller and more short-
 256 lived systems correspond to the meso- β scale (20 to 200km) according to the definition
 257 in Orlanski (1975), whereas the tracking methods described in the previous paragraphs
 258 were designed to identify larger and longer-lived MCS at the meso- α scale (200 to 2000km).
 259 We applied the same cold core and heavy rain criteria to this meso- β tracking method
 260 as for the meso- α tracking methods (Table 1) and found that reducing the size require-
 261 ment for the heavy rainfall area did not result in more systems being identified (figure
 262 not shown). All systems identified using the TCS tracking method developed a heavy rain
 263 core once during their lifetime with precipitation of at least 5 mm h^{-1} over an area of
 264 at least 5 grid cells and of at least 3 mm h^{-1} over at least 10 grid cells. The purpose of
 265 these additional runs is to investigate the role of convective cells at the lower bounds of
 266 the mesoscale over the TP. Due to limited computer capacity, the meso- β tracking could
 267 not be implemented for the entire study area as it would result in too many cloud fea-
 268 ture combinations that would have to be assessed to determine linkages across time steps.

269 Figure 4 shows the main characteristics of MCSs identified by the four different track-
 270 ing methods (T_b , T_b cold core, T_b heavy rain core, $Precip$). The diurnal cycle for fea-
 271 tures tracked using precipitation ($Precip$) has a bimodal distribution (Fig. 4a), whereas
 272 the other tracking runs are marked by a clear evening peak. The different diurnal curves

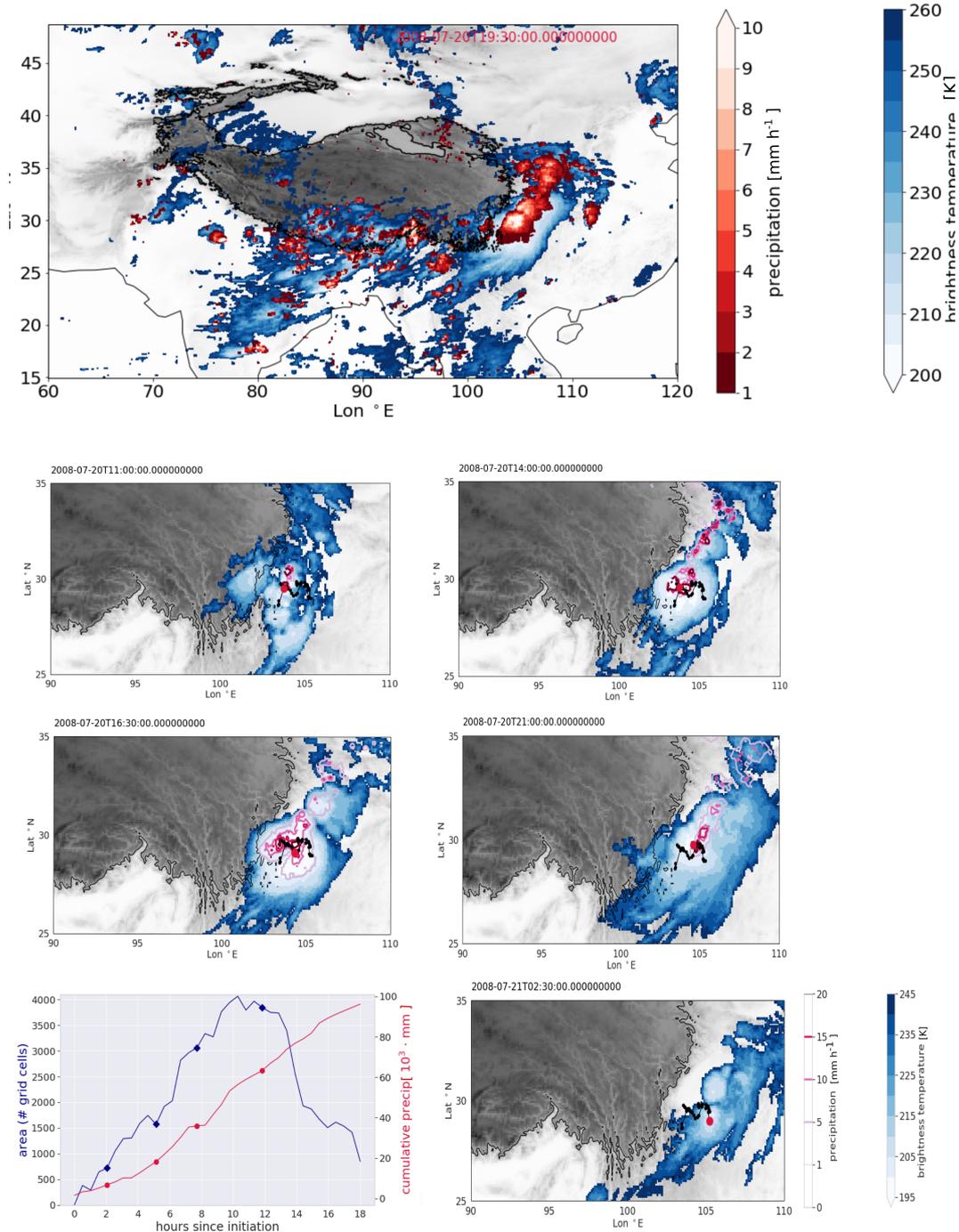


Figure 3. Example of a tracked MCS at the eastern boundary of the TP in July, 2008. The upper panel shows a snapshot of half-hourly brightness temperatures with GPM IMERG precipitation. The evolution of the tracked cloud cell and of the corresponding precipitation feature are shown in the lower panels, where the black line indicates the MCS center at the preceding and succeeding time steps and the red dot marks the center in the imaged time step. The zoomed images correspond to the time steps indicated by markers in the timeseries graph for cloud area and accumulated precipitation.

273 can be explained by the fact that *Precip* only contains the MCS features when it is rain-
 274 ing, while the other tracking methods capture the evolution of the MCSs more completely,
 275 including non-precipitating hours.

276 The T_b tracking identified a significantly higher numbers of TPB systems during
 277 winter (Fig. 4b) and generally detected more TPB systems than LE systems (Fig. 2).
 278 A similar result was found by L. Hu et al. (2017), who used MCS data from the global
 279 *ISCCP Convective Tracking Database* (C. Wang et al., 2018). By filtering with a thresh-
 280 old for optical depth, they found that the winter maximum for MCS events changed to
 281 a summer maximum. This is consistent with the well established understanding of sum-
 282 mer convection over the TP (Flohn & Reiter, 1968; Ye & Wu, 1998) and confirms the
 283 impact of MCS misclassification on climatologies, due to cold surfaces or cirrus clouds,
 284 that was discussed earlier.

285 When identified and tracked using precipitation only (*Precip*), MCSs were perceived
 286 to exhibit shorter lifetimes (Fig. 4c) and a reduced extent, which barely overlap with the
 287 MCS area distributions derived from the other three tracking methods (Fig. 4d). This
 288 is due to the fact that the precipitating area covers only a part of the cloud system, which
 289 often seems to be confined to the area with the coldest cloud top temperatures (Fig. 3).
 290 On top of that, precipitation is not necessarily contiguous in time and space, but can oc-
 291 cur in a more scattered manner when MCS are weakened and re-initiated. This also ex-
 292 plains the higher number of systems from *Precip* compared to T_b , $T_b c$ cold core and T_b
 293 heavy rain core (Table 2). As mentioned earlier, a further advantage of the combined
 294 brightness temperature-precipitation tracking (T_b heavy rain core) is, that it can cap-
 295 ture both a larger portion of the cloud cell evolution as well as the evolution of the pre-
 296 cipitating area of a cloud system.

297 As described in section 2.1, the number of TPB tracks in Table 2 contain both sys-
 298 tems which only partly cross the 3000 m boundary and systems which are entirely within
 299 the TP. Of these MCS tracks, the total number of MCSs that formed (or were first de-
 300 tected) over the TP was only 58 (around 4 MCSs per year). In contrast, tracking smaller-
 301 scale convective systems (*TCS*) over the TP resulted in identification of a total of 1428
 302 cases (around 75 cases per year). The additional systems that were identified and tracked
 303 using the *TCS* method shows the relative prevalence of smaller-scale convective systems
 304 over the TP.

305 **2.3 Analysis of MCS and associated precipitation features**

306 Since the python package *tobac* allows for feature tracking using multiple thresh-
 307 olds, each identified cloud feature is assigned to an intensity category. The intensity cat-
 308 egories are defined as areas within the detected cloud feature where a specific brightness
 309 temperature threshold is exceeded (between 195 K and 230 K). We use the cloud fea-
 310 ture characteristics at each time step (shape, area, brightness temperature intensity, pre-
 311 cipitation features) and the characteristics of the track that describes the MCS evolu-
 312 tion (lifetime, total precipitation, propagation speed, propagation direction) to compare
 313 different MCS types. Firstly, all tracked MCSs in the study area (Fig. 3) were assigned
 314 to one of the four classes based on their propagation direction (northward and eastward)
 315 and location (denoted as TPB east, TPB north, LE east and LE north). These two prop-
 316 agation directions were chosen, because most of the detected MCS trajectories follow one
 317 of these directions, which match the directions of major large-scale atmospheric circu-
 318 lation systems that affect the region (e.g. mid-latitude westerlies and the northward prop-
 319 agating Indian summer monsoon). The total amount of eastward-moving TPB systems
 320 were slightly higher than for LE features, but the northward-moving LE systems exceeded
 321 the number of northward-moving TPB systems.

322 To investigate the importance of MCS for precipitation from a climatological per-
 323 spective, the total amount of precipitation is calculated for each of the detected cloud

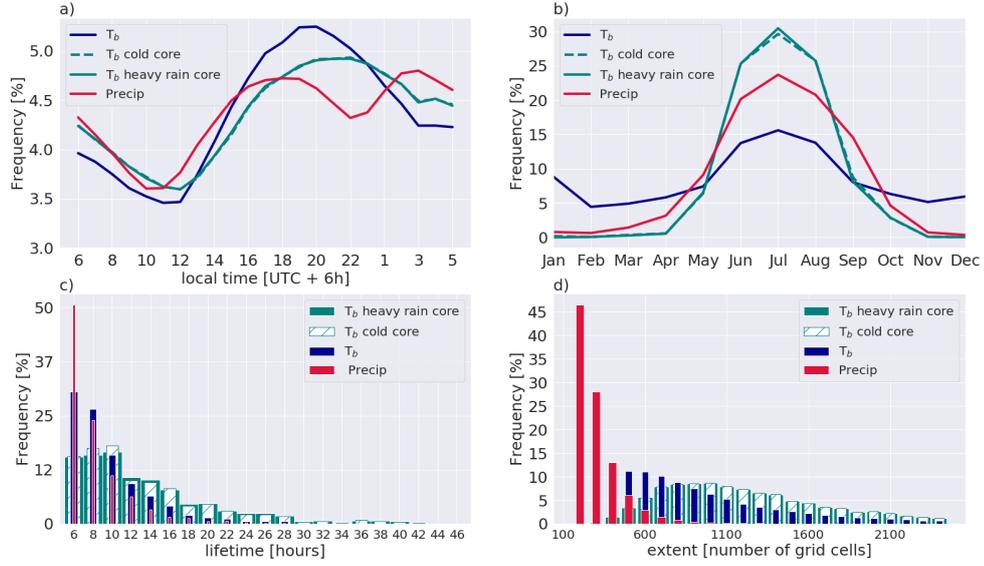


Figure 4. Comparison of characteristics for MCSs identified using four different tracking methods. The subfigures show the a) average diurnal cycle for the frequency with which a feature was identified within the entire study area b) average seasonal distribution for features identified over the TP c) distribution of the lifetime [h] for all MCS tracks and d) distribution of the extent [number of grid cells] of all tracked MCS features.

Table 2. Total number of MCS between 2000 and 2019 using different tracking methods.

| MCS type | Tracking method | number of tracks (average per year) |
|----------|-----------------------|--|
| LE | <i>Precip</i> | 5465 |
| TPB | <i>Precip</i> | 4090 |
| total | <i>Precip</i> | 9555 |
| LE | T_b | 563 |
| TPB | T_b | 3117 |
| total | T_b | 3680 |
| LE | T_b cold core | 294 |
| TPB | T_b cold core | 261 |
| total | T_b cold core | 555 |
| LE | T_b heavy rain core | 277 |
| TPB | T_b heavy rain core | 240 |
| total | T_b heavy rain core | 517 |
| TP | <i>TCS</i> | 75 |

324 features and compared to total monthly and seasonal precipitation received at every grid
 325 cell. Other precipitation mechanisms than MCSs may dominate the total annual and sea-
 326 sonal precipitation, and so we also examine the importance of MCS-associated rainfall
 327 only for extreme precipitation events. The same reasons that were given for the choice
 328 of the rain rate threshold in the heavy rain core of an MCS (see section 2.2), motivated
 329 us to define extreme convective precipitation as all precipitation that falls at rain rates
 330 equal to or higher than 5 mm h^{-1} .

331 3 Results

332 3.1 Spatial and temporal characteristics

333 Figure 5 shows histograms of the seasonal occurrence (a), mean propagation speed
 334 (b), lifetime (c) and mean extent (d) for LE north, LE east, TPB north and TPB east.
 335 The seasonal cycles for TPB systems have a distinct peak in June (Fig. 5a), which is di-
 336 rectly follows the onset of the Indian summer monsoon season over the TP. Both LE east
 337 and LE north also exhibit an early summer peak of MCS occurrences, which is, however,
 338 much more pronounced than MCSs in the TPB region. While more than 75 % of TPB
 339 east and TPB north occur between June and October, the frequency of MCSs that oc-
 340 cur in LE north and LE east is significantly smaller during autumn and drops by around
 341 10 % from the month of September.

342 The distributions of mean propagation speed, lifetime and mean extent are sim-
 343 ilar for the four MCS types LE north, LE east, TPB north and TPB east (Fig. 5b-d),
 344 showing that MCS in each subgroup still exhibit a large variety of characteristics that
 345 are not primarily affected by their propagation direction (mean environmental wind) and
 346 elevation. More than half of the tracked MCS in each subgroup do not last longer than
 347 12 hours, but all four MCS types include examples of long-lived MCS that last longer
 348 than 30 hours (Fig. 5c). The distribution of MCS extent is right-skewed, similar to the
 349 distribution of MCS lifetime, and MCSs with the greatest extent are slightly more fre-
 350 quent for LE north and LE east types than for TPB east and TPB north types (Fig. 5d).
 351 Most of the tracked MCSs have a mean extent between 1000 and 1500 grid cells, which
 352 corresponds to an area of about $100\,000 - 150\,000 \text{ km}^2$. This shows that the area of the
 353 cloud shield for most of the systems is at least 10 times as large as the required mini-
 354 mum cold area (Fig. 5d) and that the horizontal dimensions of of the dominating MCSs
 355 type in the study area are comparable to *Mesoscale Convective Clusters* (described in
 356 section 2).

357 It can be seen clearly in Figure 6 that the Indian subcontinent and adjacent ocean
 358 are the regions, where most of the MCSs originate. The Indo-Gangetic Plain and the Hi-
 359 malayas are the initiation location for TPB systems, rather than the TP itself (Fig. 6a-
 360 b). The few MCSs that initiate (or first detected) over the TP mainly form over the east-
 361 ern parts of the TP and are, except for a few cases, transported eastward (Fig. 6a). MCSs
 362 that are transported towards the TP and are therefore included in the TPB class may
 363 have travelled large distances, but they do not reach far behind the Himalaya mountain
 364 ranges, but stop south of the Himalayas, where most of the TP rainfall occurs (Kukulies
 365 et al., 2020).

366 To compare MCSs over the TP with systems in the surrounding regions, we selected
 367 TPB systems that formed over the TP and either remained within the boundary of the
 368 TP or moved out of the TP. Figure 3.1a) shows the location of the centers of these sys-
 369 tems at the time of initiation (when they first tracked as a cloud feature), when they reach
 370 maturity (the age of a cloud feature when its embedded area of precipitation with at least
 371 5 mm h^{-1} is greatest) and when they dissipate (the last time of they area tracked as a
 372 cloud feature). It is important to understand the diurnal evolution of MCSs that origi-
 373 nate over the TP, because they may be closely linked to topographically-driven diur-

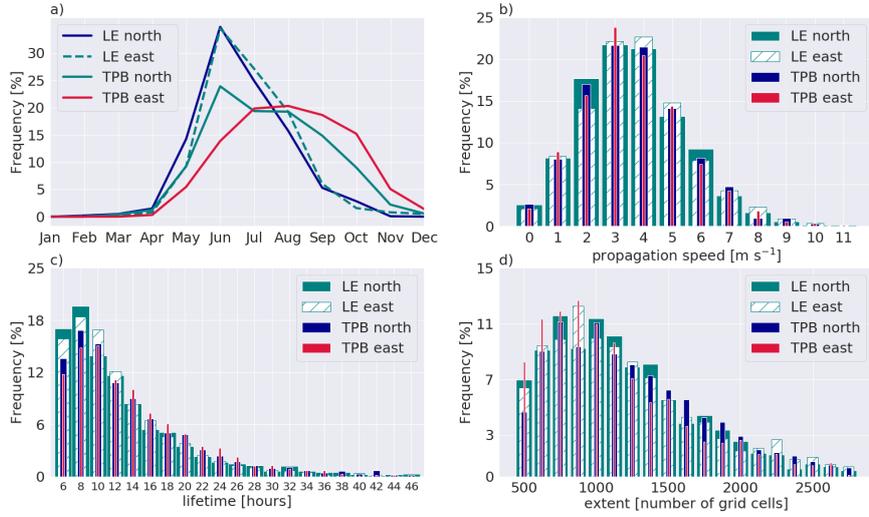


Figure 5. Spatial and temporal characteristics of northward- and eastward-moving MCSs in the LE and TPB regions. The histograms show the relative frequencies [%] for a) seasonal occurrence [%], b) mean propagation speed [m s^{-1}], c) lifetime [h] and d) mean extent [number of grid cells].

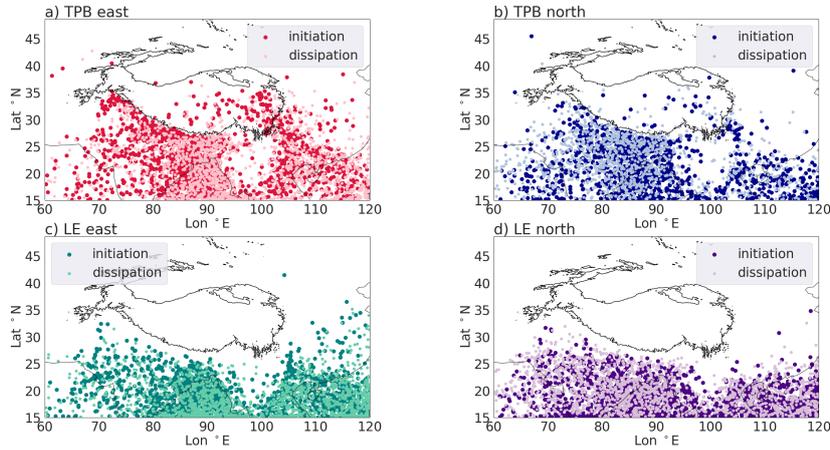


Figure 6. Location of MCS centers at the time of initiation and dissipation for a) TPB east, b) TPB north, c) LE east and d) LE north. Markers depict the mean longitude and latitude for detected cloud features at initiation (the time step at which it is first detected) and dissipation (the last time step in which the feature is detected).

374 nal flow patterns. Figure 3.1b) suggests that MCSs follow a common pattern of diurnal
 375 evolution over the TP, with initiation around noon (12:00 LST), maturity and dissipa-
 376 tion occurring late night or in the early morning (between 00:00 and 5:00 LST). The re-
 377 sulting evening and night peaks in precipitation are consistent with the general diurnal
 378 cycle of precipitation over the TP (Kukulies et al., 2019) and indicates that the diurnal
 379 flow is an important component for the organization of convective cells into larger sys-
 380 tems.

381 **3.2 Precipitation characteristics**

382 **3.2.1 Contribution to total and extreme precipitation**

383 Figure 8 shows seasonal contributions of precipitation from tracked MCS cloud fea-
 384 tures to the total precipitation and to the total extreme precipitation (where extreme
 385 precipitation is from rainfall events with a rate of at least 5 mm h^{-1}). During spring,
 386 the highest MCS contribution to rainfall is over the ocean (Bay of Bengal), and the spa-
 387 tial pattern of the contribution is similar for both total and total extreme precipitation
 388 (Fig. 8a-b). Locally, rainfall from MCSs accounts for up to 30 % of the total rainfall over
 389 land, but less than 1% of the precipitation over the TP (Fig. 8a-b). Between June and
 390 August, the contribution of MCSs to both total and total extreme precipitation increases
 391 substantially (Fig. 8c-d). Although the MCS contributions to precipitation over the TP
 392 are significantly lower than the contributions for the surrounding regions, the con-
 393 tributions to extreme precipitation are between 30 and 70 % to the south of the Himalayas
 394 and over the Indo-Gangetic Plain. The largest contributions over the TP are in the east-
 395 ern part of the TP, where MCS precipitation accounts for up to 30 % of the seasonal lo-
 396 cal rainfall. MCS precipitation contributes significantly to the seasonal rainfall over land
 397 outside the TP (Fig. 8c-d), and there are only few local maxima over the ocean and at
 398 the coast in India for the contribution in winter (Fig. 8g-h).

399 These results confirm our hypothesis that organization of convective systems, which
 400 contribute to total summer precipitation and extreme rainfall over the TP occurs on a
 401 smaller scale here than in the surrounding regions, where MCSs provide a large portion
 402 of both the total extreme rainfall and the total summer rainfall. In comparison, the sys-
 403 tems tracked over the TP (*TCS*) using a smaller area threshold (see section 2.2 and Ta-
 404 ble 1) make a significantly higher contribution to precipitation locally during the sum-
 405 mer months (Fig. 9). In contrast to the June peak of MCSs for TPB east and TPB north,
 406 most of the identified smaller-scale convective systems over the TP occurred between July
 407 and August. These systems account for between 25 and 35 % of the total monthly rain-
 408 fall over large areas of the central and eastern TP. We also see a strong difference be-
 409 tween the contribution of these systems to total and to total extreme precipitation. The
 410 contributions to extreme precipitation have strong local maxima and therefore a patchy
 411 spatial pattern with many grid cells for which more than 50 % of local extreme precip-
 412 itation is accounted for by tracked systems. A large area of high values for the contri-
 413 bution occurs at the eastern edge of the TP (Fig. 9). This region is the same region that
 414 had the highest values MCS contributions over the TP (Fig. 8). There is no clear pat-
 415 tern linking the contribution to precipitation with topography, but it should be noted
 416 that small-scale convective systems contribute to summer precipitation even at eleva-
 417 tions greater than 5000 m a.s.l. (Fig. 9). This means that organization of convection over
 418 a few 10^1 km is not confined to the lower elevations over the TP.

419 **3.2.2 Convective precipitation**

420 The snapshot in Figure 3 is a typical example of an MCS, and shows that the con-
 421 vective core with heavy precipitation is often surrounded by a larger area of stratiform
 422 precipitation at moderate rain rates and by an even larger area of non-precipitating clouds.
 423 Although these stratiform cloud shields may extent over a greater area, the convective

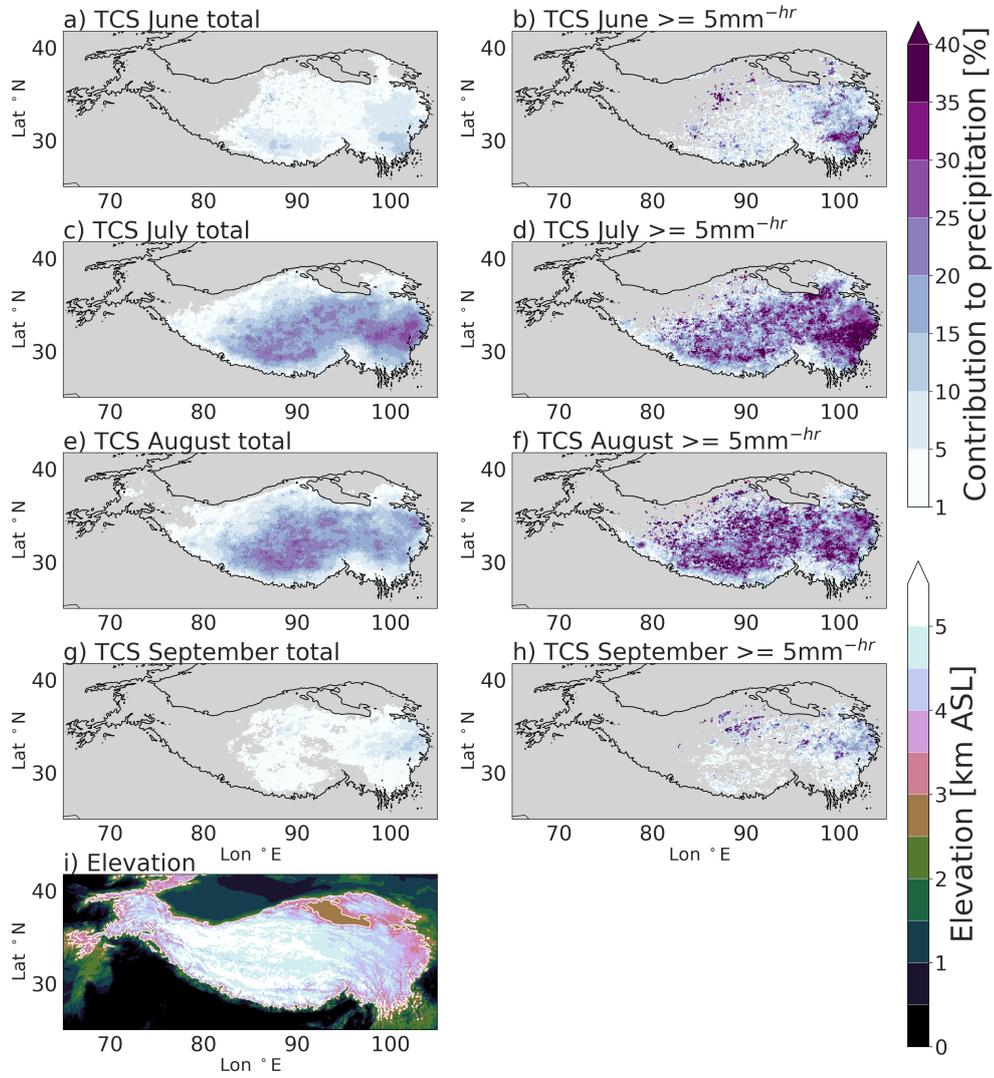


Figure 7. Spatial and temporal evolution of MCSs that initiated over the TP between 2000 and 2019 (58 cases). The left panel (a) shows locations for the MCS centers at initiation, maturity (time associated with maximum precipitation) and dissipation. The right panel (b) shows the average time of the day for initiation, maturity and dissipation of MCSs over the TP.

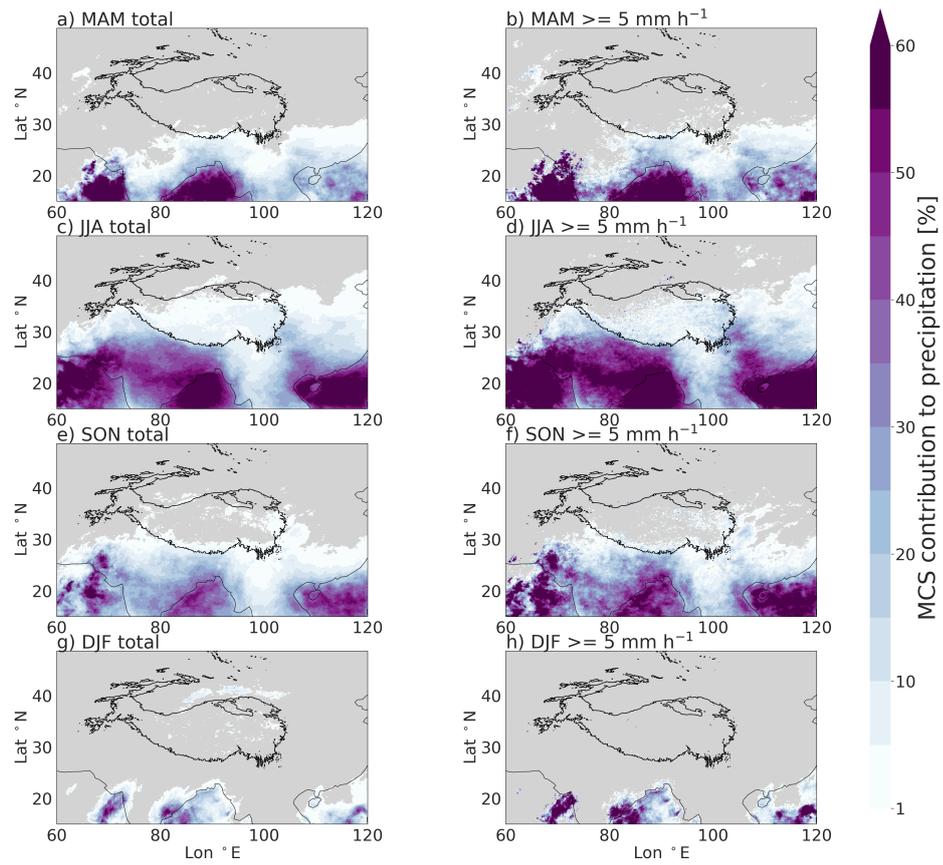


Figure 8. Maps of the seasonal contribution of precipitation from MCS [%] to the total precipitation (a,c,e,g) and to the total extreme precipitation, which is the sum of precipitation produced by rainfall of at least $\geq 5 \text{ mm h}^{-1}$ (b,d,f,h).

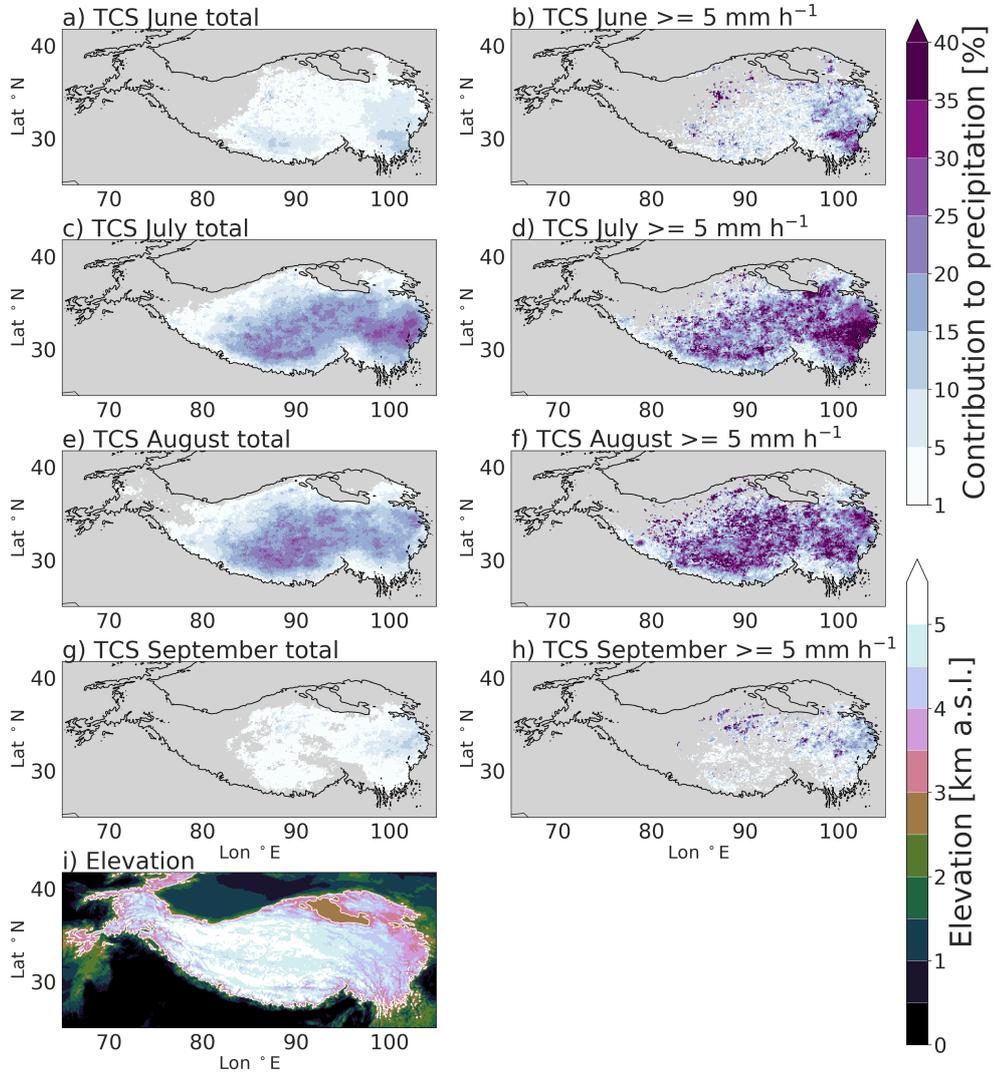


Figure 9. Maps of the contributions to total and total extreme precipitation from smaller and more short-lived convective cells over the TP (*TCS*) [%]. These convective cells have approximately meso- β dimensions in the horizontal plane (20 to 200 km). Contributions are shown for June to September as the ratio of precipitation from the convective cells to total precipitation (a,c,e,g) and to total extreme precipitation (b,d,f,h), defined as the sum of precipitation that fell at a rate of at least 5 mm h^{-1} . The elevation of the TP [km a.s.l.] is shown for context (i).

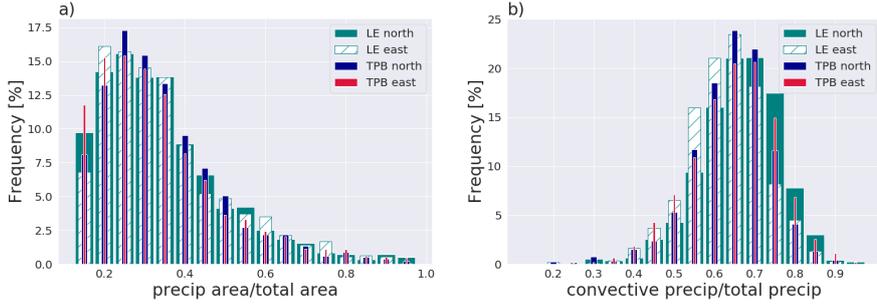


Figure 10. Histograms of a) the fraction of total cloud area that precipitates [precipitating area/total area] and b) the fraction of the precipitating cloud area associated with convective precipitation [convective precipitating/total precipitating area]. Convective precipitation is here defined as precipitation at a rate of at least 5 mm h⁻¹.

424 part of an MCS produces large amounts of rainfall over short periods and is therefore
 425 of higher environmental and hydrological relevance than the stratiform part. The dis-
 426 tribution of the precipitating area, expressed as a fraction of the total cloud area for all
 427 identified MCSs, is right-skewed, which shows that for more than half of the identified
 428 MCSs, only 10 % to 40 % of the detected cloud area produces precipitation (Fig. 10a).
 429 Only a small part of the entire cloud complex of an MCS produces precipitation, and
 430 the convective core usually covers an area even smaller than this. For more than 90 %
 431 of all identified and tracked MCSs, the histogram of the fraction of the precipitating area
 432 that corresponds to convective precipitation peaks between 0.6 and 0.8, which means that
 433 60 % to 80 % of all precipitation from MCSs falls at rates of at least ≥ 5 mm h⁻¹ (Fig.
 434 10b). This is consistent with other studies, which have shown that around 40 % of pre-
 435 cipitation in well-developed MCSs is stratiform, while around 60 % of the total rainfall
 436 is of convective type (Cheng & Houze, 1979; Rutledge & Houze Jr, 1987). LE north sys-
 437 tems have generally larger convective contributions to total rainfall than the other three
 438 MCS types and TPB east systems have the smallest convective precipitation contribu-
 439 tions (Fig. 10b).

440 Results from the different MCS tracking methods were similar, regardless of whether
 441 the cold core or the heavy rain core criterion was applied, as mentioned in section 2.2
 442 (Table 1). This implies that MCS precipitation intensity is reflected in the observed bright-
 443 ness temperatures, with cooler temperatures corresponding to greater intensity. Figure
 444 11) shows the average rain rates for the different MCS types as a function of the low-
 445 est brightness temperature threshold above which a contiguous area of at least 100 pix-
 446 els exist. It is shown that MCS that have lower brightness temperatures also have higher
 447 average rain rates (Fig. 11). This suggests that more intense precipitation is produced
 448 by deeper cloud systems, which have higher cloud tops and therefore correspond to lower
 449 brightness temperature in IR images.

450 3.2.3 Heavy impact MCS

451 The total rainfall amount produced by an MCS depends on the system’s lifetime,
 452 size and intensity. The four MCS types presented in Figure 5 include MCSs with highly
 453 variable track characteristics, and so we further divide each MCS type into three cat-
 454 egories, according to lifetime, size and intensity (Fig. 12). To identify what character-
 455 izes an MCS with a large environmental impact, the different MCS classes are compared
 456 based on their total amount of convective precipitation they produce during their life-
 457 time. Figure 12 shows the distributions of total convective precipitation for the three cat-

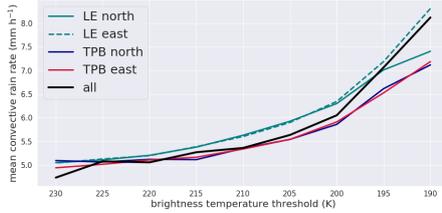


Figure 11. Mean convective rain rate [mm h^{-1}] as a function of the brightness temperature [K] for the coldest contiguous area in one MCS.

458 categories of lifetime, size and intensity for the MCS types LE north, LE east, TPB north
 459 and TPB east. The boxplot shows that the total amount of convective precipitation varies
 460 substantially between the tracked MCS. Differences between the total convective pre-
 461 cipitation distributions for the four previously defined MCS types (LE north, LE east,
 462 TPB north, TPB east) are small, which supports our earlier finding that large-scale at-
 463 mospheric systems, which determine MCS transport and genesis location, do not deter-
 464 mine the environmental impact of the systems. Hence, both eastward- and northward-
 465 moving MCSs over lower and higher elevations, may produce substantial amounts of heavy
 466 rainfall. Comparing lifetime, area and intensity, for systems in the different categories,
 467 the most obvious difference is between systems that last longer than 24 hours and shorter-
 468 lived systems (Fig. 12a). The mean, maximum and outliers for the distributions of to-
 469 tal convective precipitation all increase as the area covered by the system increases (Fig.
 470 12b), and the total convective precipitation is not obviously related to the temperature
 471 of the convective part of the system (Fig. 12c).

472 MCSs which initiate over the TP (Fig. 3.1) are marked in red (Fig. 12) and are
 473 located in the lower-precipitation part of the distribution for each category. This means
 474 that systems which initiate over the TP produce little convective precipitation compared
 475 to those that form outside, or at the edges of the TP. The MCS observed on July 2008
 476 (Fig. 3) is also marked in Figure 12 for context (red cross). This MCS persisted for 18
 477 hours and produced more convective precipitation than most systems over the TP. It falls
 478 into the largest mean category, but produced less total convective precipitation than other
 479 systems which persisted for more than 24 hours (Fig. 12a-b).

480 The joint frequency distributions for total convective precipitation and maximum
 481 rain rate, system lifetime, maximum convective area and mean temperature of the low-
 482 est brightness temperature for a contiguous area (Fig. 13), show that lifetime and area
 483 have a greater effect on the environmental impact of the systems than rain rate inten-
 484 sity or brightness temperatures.

485 As stated earlier, MCSs may take different geometric shapes, ranging from elon-
 486 gated squall lines to quasi-circular MCCs. More elongated, or linear-shaped, systems have
 487 an eccentricity [minor axis/major axis] closer to 0, whereas more circular-shaped MCS
 488 have an eccentricity closer to 1. Figure 3.2.3a shows an example snapshot of a tracked
 489 MCS (August 17th, 2012 at 13:30 UTC) with an eccentricity value of 0.77, which is typ-
 490 ical value for an MCC (Maddox, 1980). As shown in Figure 3.2.3b, most systems tracked
 491 in the TP region have oval or ellipsoid shapes with eccentricity values between 0.3 and
 492 0.6. We define MCSs with the highest environmental impact as those in the 90th per-
 493 centile of the total convective precipitation-distribution. These systems tend to be slightly
 494 more circular-shaped, relative to the eccentricity distribution that includes all tracked
 495 MCSs (Fig. 3.2.3b). Figure 3.2.3c,d shows the eccentricity distributions divided into the
 496 four subgroups LE north, LE east, TPB north and TPB east. All tracked systems in the

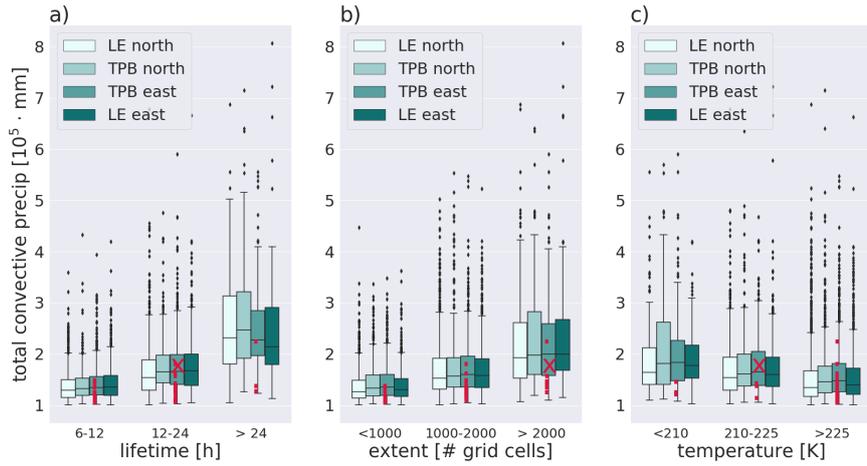


Figure 12. Boxplot showing the distribution of total convective precipitation produced by eastward- and northward-moving MCSs which cross the TP (TPB) and which remain in the surrounding lower-elevation plains (LE). MCSs are divided into different classes dependent on a) lifetime [h], b) extent [number of grid cells] and c) temperature [K]. The temperature refers to the lowest brightness temperature for a contiguous area within the cloud feature. The red dots show systems within the TPB group that remained over the TP and the red cross highlights the MCS from of July 2008 used as a case study (see section 2).

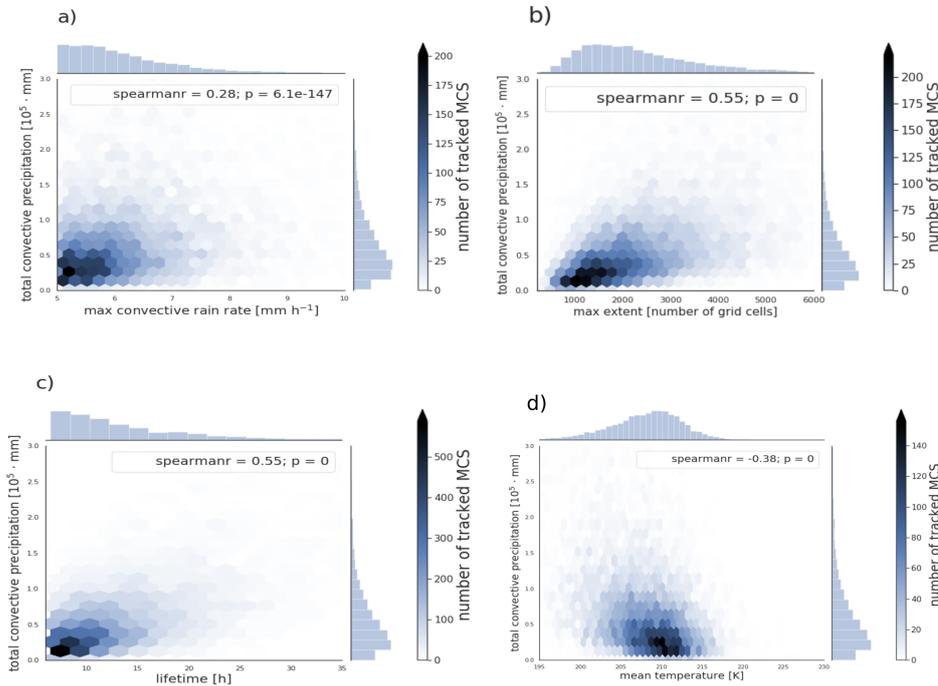


Figure 13. Frequency density plots for the total convective precipitation and a) maximum convective rain rate [mm h^{-1}], b) maximum convective area [number of grid cells] and c) lifetime [h] and d) mean brightness temperature [K] for all tracked MCS. The color scale reflects the number of identified MCS tracks which correspond to each part of the plotted space.

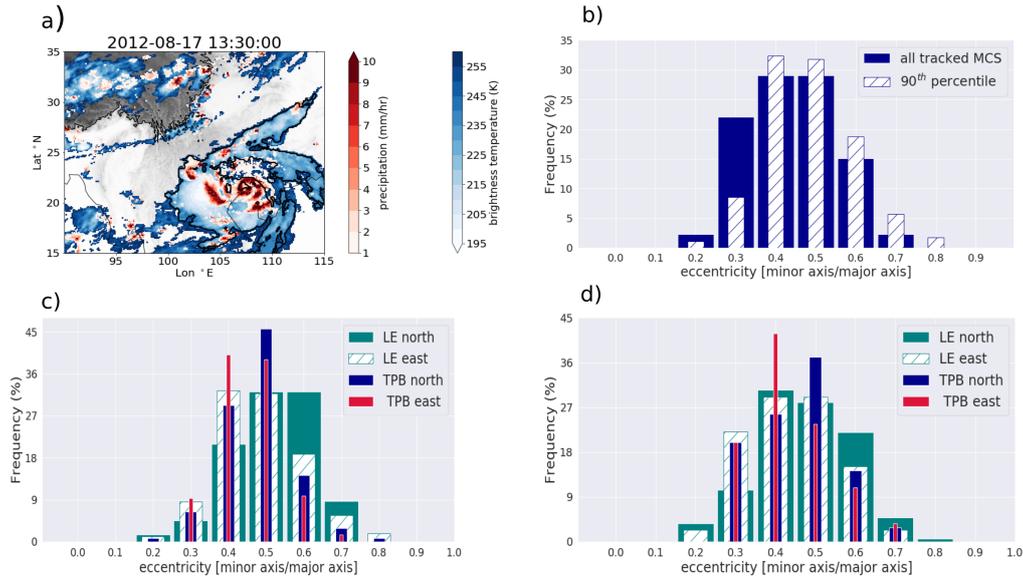


Figure 14. Eccentricity [minor axis length/major axis length] for tracked MCSs. a) An example of a quasi-circular MCS (August 17th, 2012 at 13:30 UTC) with an eccentricity of 0.77, b) Histogram of eccentricity values for all tracked MCSs and for MCSs which produce the largest total amount of convective precipitation (rain rate of at least ≥ 5 mm h^{rr}). These systems make up the 90th percentile of the distribution of total convective precipitation; c) Histogram of eccentricity values for systems, divided into MCS classes: LE north, LE east, TPB north, TPB east; and d) same as c) but shown separately for MCS classes: LE north, LE east, TPB north, TPB east.

497 respective class are included in the distributions in (c), and the distributions in (d) include
 498 only systems that have the highest environmental impact, i.e. from the 90th percentile of the total
 499 convective precipitation distribution for that class. It is clear that LE north and LE east systems are
 500 generally more circular-shaped than TPB east and TPB north systems (Fig. 3.2.3c), and MCSs with
 501 eccentricity greater than 0.5 become more frequent for heavy impact MCSs in the TPB east and
 502 TPB north classes (Fig. 3.2.3d). The figure shows that squall lines, which have a great environmental
 503 impact over the US Great Plains (Yang et al., 2017; Hitchcock et al., 2019), are not present in our study
 504 area. In contrast to the Great Plains in the USA, the most intense systems in our study area
 505 have slightly higher eccentricity values, particularly when the MCS is close to the TP,
 506 where the terrain most likely inhibits any quasi-linear organization of convection.
 507

508 4 Discussion

509 4.1 Role of MCS for precipitation

510 Previous studies have highlighted the importance of MCS-induced precipitation as
 511 being the main source of summer precipitation over the TP. The reason for the discrep-

512 ancy between that conclusion and our findings is most two-fold. Our method for MCS
 513 tracking is less likely to include cirrus clouds than more standard tracking methods, and
 514 our method for calculating MCS-associated precipitation differs from the commonly used
 515 methods. Many studies use a radius approach (L. Hu et al., 2017; Curio et al., 2019),
 516 where all precipitation within a certain radius of the MCS center is considered to be MCS-
 517 induced precipitation, instead of tracking precipitation features within cloud cells as we
 518 have done here. Our results show that the precipitating area of an MCS can vary sig-
 519 nificantly (Fig. 10), and the method we have used here therefore gives a more accurate
 520 estimation of precipitation associated with an MCS, relative to the size of the MCS. More-
 521 over, we only track larger MCSs and exclude systems at the meso- β scale, which, as we
 522 could show, have higher contributions to summer precipitation over the TP.

523 This study shows that isolated, smaller-scale (meso- β) convective cells make a greater
 524 contribution than MCSs (meso- α) to both seasonal mean precipitation and extreme sum-
 525 mer precipitation, according to our tracking method. This result suggests that other con-
 526 vective modes, which are not captured by the applied tracking methods, are important
 527 and that topography and moisture conditions over the TP may inhibit the organization
 528 of convective cells into large systems. A similar conclusion was drawn by Houze et al.
 529 (2007) who investigated deep convective features based on Tropical Rainfall Measuring
 530 Mission (TRMM) Precipitation Radar (PR) data and found that deep convective echoes
 531 occur over the TP in a scattered manner, whereas more prominent mesoscale convective
 532 features organized along the Himalayan ranges. More complex convective modes (e.g.
 533 multi-cellular convective cells), which are not captured by the applied tracking, and
 534 convective cells with smaller horizontal extents are most likely more important for extreme
 535 precipitation. A possible explanation for the presence of isolated convective cells could
 536 be the occurrence of cirrus clouds at higher altitudes, which may locally inhibit convec-
 537 tive heating during summer (Roebber et al., 2002). Satellite observations reveal a bimodal
 538 distribution of cloud top heights over the TP (D. Chen et al., 2018; Kukulies et al., 2019),
 539 whereby cloud top temperatures below 220 K correspond to cloud top heights above 10
 540 km a.s.l. (D. Chen et al., 2018). The presence of isolated convective cells could there-
 541 fore be an indicator of simultaneous occurrence of cirrus clouds and convective clouds.

542 However, occasionally larger MCSs according to the definition used here were also
 543 found over the TP (in total 58 cases). Most of the MCSs that formed over the TP were
 544 eastward-moving and resulted in heavy precipitation at the eastern edges of the TP. Al-
 545 though the amount of convective precipitation was relatively small compared to that from
 546 MCSs over the Indo-Gangetic Plain (due to generally lower intensity and lifetimes), these
 547 MCSs can be very destructive when they enter a populated downstream region and cause
 548 devastating floods, such occurred in July 2008 (X. Feng et al., 2014). More studies are
 549 needed to investigate future projections of both MCS intensity, but also MCS frequency
 550 over the TP and in the surrounding regions.

551 In the LE and in parts of the TPB region, the importance of MCSs for summer mean
 552 and extreme rainfall in summer was significantly higher than over the TP. Densely pop-
 553 ulated regions south of the Himalayas and in the Indo-Gangetic Plain experience frequent
 554 MCS events, which produce substantial rainfall amounts due to their longevity and large
 555 size. Z. Feng et al. (2018) found that long-lived MCS over the Great Plains in the USA
 556 produced 2–3 times more precipitation than short-lived MCSs. A similar result was found
 557 in this study, where both mean and maximum total convective precipitation of MCSs
 558 that persisted longer than 24 hours were twice as high as MCSs that persisted between
 559 12 and 24 hours and thrice as high as MCS that persisted for up to 12 hours. The fact
 560 that there were no large differences in the contribution of precipitation from MCSs to
 561 the total seasonal and total extreme precipitation, shows that MCSs south of the TP are
 562 not only important for local extremes, but also important over climatological timescales.
 563 Changes in MCS patterns do not only affect the risk of severe weather for populated re-
 564 gions, but can also lead to changes in the hydrological cycle. It should be noted that the

565 contribution of MCSs to precipitation may be underestimated for these regions in this
 566 study, because our MCS tracking method was optimized for the TP. Since cloud top tem-
 567 peratures for deep convective cells of the same depth are higher over lower-elevation re-
 568 gions than over regions at higher elevation, the differences between MCS-induced pre-
 569 cipitation over the TP and in the LE region (such as the Indo-Gangetic Plain) may be
 570 even larger than is shown in our results.

571 To get a more complete understanding of the convective modes and their precipi-
 572 tation structure over the TP, further studies should consider the vertical aspects of the
 573 organization of convection. Radar observations of clouds over the southern Himalayas
 574 show mean radar reflectivities with convective signals over a height range between 3 and
 575 14 km a.s.l. during summer (Kukulies et al., 2019). This may indicate MCSs in that re-
 576 gion, but the temporal poor resolution of spaceborne radar observations means that there
 577 is a need for ground radar observations to confirm and investigate this.

578 4.2 Possible mechanisms for MCS formation

579 This study provides an observational perspective on MCSs over a region, where multi-
 580 ple processes may lead to the organization of convection at the mesoscale. Most of MCSs
 581 in the LE region over land were found south of the TP, over the Indo-Gangetic Plain and
 582 close to the Himalayas, as well as over the southern Indian subcontinent and the Bay
 583 of Bengal. Both land and ocean south of the TP are dominated by monsoon low pres-
 584 sure systems during the wet season, and these have been associated with barotropic in-
 585 stabilities and that may be amplified by wind-moisture feedbacks (Boos et al., 2017; Diaz
 586 & Boos, 2019). The results of this study show clearly that monsoon low pressure sys-
 587 tems appear as large cloud clusters with similar dimensions as MCC ($\sim 100\ 000\ \text{km}^2$).

588 The mechanisms for organization of convection over the TP are likely to be differ-
 589 ent because of the relatively limited moisture supply and the influence of local wind sys-
 590 tems. The distinct diurnal evolution of MCSs (Fig. 3.1b), that are initiated over the TP
 591 and the large contributions of smaller-scale convective cells to precipitation suggest the
 592 importance of local conditions and topography. Surface properties, such as soil moisture
 593 have been suggested to be major contributing factor for MCS formation for the moist
 594 regions of the eastern TP (Sugimoto & Ueno, 2012). The systems that made the great-
 595 est convective contributions to precipitation over the TP in this study (including both
 596 MCSs and the smaller-scale convective cells) were also located in the east. The eastern
 597 TP is characterized by lower elevations and higher soil moisture, which may be crucial
 598 factors for the organization of convection. Outside of the monsoon season, when mois-
 599 ture supply to the TP is more limited, MCS could possibly form due to the lake-effect,
 600 which has been shown to trigger severe storms during winter (Dai et al., 2020).

601 While baroclinic instabilities are not among the the most important forcings for
 602 monsoon depressions (Diaz & Boos, 2019), upper-level baroclinicity has been found to
 603 be correlated with both different mesoscale disturbances over the TP, namely Tibetan
 604 Plateau vortices (TPV) (X. Feng et al., 2017; Curio et al., 2019), and Western distur-
 605 bances (Hunt et al., 2018). This effect is mainly controlled by the strength and position
 606 of subtropical westerly jets when these systems form (Hunt et al., 2018). Mesoscale dis-
 607 turbances in relative vorticity occur frequently over the TP at around 500 hPa, particu-
 608 larly during the summer season (Curio et al., 2019). Thus, it should be noted that scat-
 609 tered precipitation features in satellite observations over the TP do not imply that mesoscale
 610 forcings are irrelevant for precipitation formation. This was demonstrated by Curio et
 611 al. (2019), who showed that a significant part of the plateau-scale precipitation occurs
 612 within a 3° radius of tracked Tibetan Plateau vortex centers. Even if these precipita-
 613 tion features do not occurred as contiguous areas of precipitation or embedded in con-
 614 tiguous cloud clusters, these may be still triggered by mesoscale disturbances. Studies
 615 into the linkages between TPVs and observations of MCSs and smaller-scale convective

616 cells, could provide valuable insights into mesoscale dynamics over the TP. Fur-
 617 ther, convection-permitting simulations could provide a more complete picture of the un-
 618 derlying dynamics for precipitation formation over the TP. In other regions, such as in
 619 North America, it could be shown that convection-permitting simulations realistically
 620 capture the main characteristics of MCSs and associated precipitation (Prein, Liu, Ikeda,
 621 Bullock, et al., 2017). So far, there are no studies that have looked at MCSs over the TP
 622 using model simulations with higher than 30 km-spatial resolution, although a few sim-
 623 ulations with finer resolutions exist. The results from this study suggest that model sim-
 624 ulations at higher resolutions are necessary for the effective representation of convective
 625 features over the TP region, particularly at elevations above 3000 m, where convective
 626 cells have much smaller spatial extents, relative to those over lower elevation surfaces.
 627 The existing and future fine resolution modeling over the TP should be used to explore
 628 the dynamics of the MCSs.

629 5 Summary and conclusions

630 This study provides an observational overview of mesoscale convective systems (MCS)
 631 in the Tibetan Plateau (TP) region and elucidates the role of MCSs for seasonal and ex-
 632 treme precipitation. We tracked MCSs by combining brightness temperatures from IR
 633 satellite imagery with the precipitation product GPM IMERG for the period 2000 to 2019.
 634 Spatial and temporal characteristics of MCS tracks and the structure of MCS-induced
 635 precipitation were examined for MCSs over the TP, MCSs that cross the TP boundary
 636 (TPB) and MCS over the surrounding lower-elevation plains (LE).

637 By comparing four different MCS tracking methods for the TP region, we have shown
 638 that it is useful to apply an additional criterion for more accurate MCS identification.
 639 To be considered an MCS, a cloud feature must contain either a region below 200 K (T_b
 640 cold core) or a region with rain rates equal to or above 5 mm h^{-1} that extends over at
 641 least ten grid cells (T_b heavy rain core). This reduces the number of falsely identified
 642 MCSs due to the presence of cirrus clouds or cold surfaces in higher mountain regions,
 643 resulting in a more realistic seasonal cycle for MCS frequency with a summer peak. There
 644 is a strong relationship between low brightness temperatures and heavy precipitation,
 645 and the T_b cold core and T_b heavy rain core criteria therefore resulted in identification
 646 of almost the same number of MCS tracks and similar MCS characteristics. This implies
 647 that these criteria can be used interchangeably, but the use of both precipitation and bright-
 648 ness temperature remains advantageous because it allows precipitation features to be iden-
 649 tified and examined in the tracked clouds cells.

650 Most of the MCSs identified using our MCS tracking method were found over the
 651 Indian subcontinent and Bay of Bengal. Major regions of genesis for MCSs that crossed
 652 the TPB were the Indo-Gangetic Plain and the south of the Himalayan mountain range.
 653 In these regions, MCSs account between 10 % and 55 % of total precipitation between
 654 July and August and between 10 % and 70 % to total extreme precipitation in summer.
 655 A large part of the tracked MCS in the study area are characterized by quasi-circular
 656 cloud cells with large areas that are comparable to Mesoscale Convective Clusters (MCC).
 657 Further, MCS which produced the highest total convective precipitation amounts were
 658 characterized by longevity and large cloud extents rather than by high intensities.

659 MCSs over the TP were less frequent and represented only a small subgroup of MCSs
 660 identified at the TPB. We have shown that the contribution of MCS-induced precipita-
 661 tion to the total summer precipitation (total summer extreme precipitation) was only
 662 1 % to 10 % (1 % to 30 %), which is significantly smaller than has been found in pre-
 663 vious studies and is lower than for the surrounding regions. Additionally, the total amount
 664 of convective precipitation produced by MCS over the TP was also smaller compared to
 665 its surroundings, since MCS over the TP were generally smaller and more short-lived.
 666 In contrast, the results of tracking smaller-scale (meso- β scale) and more short-lived con-

667 vective cells showed that these made significantly higher contributions to the total and
 668 the total extreme precipitation over the TP in July and August. This result highlights
 669 the significance of more localized precipitation systems and convective modes to total
 670 summer precipitation over the TP. Many studies focus on frequently occurring mesoscale
 671 vorticity disturbances over the TP, for example Tibetan Plateau vortices (TPV) and West-
 672 ern Disturbances, and the results from this study suggest that there is a more complex
 673 relationship between mesoscale dynamics and the scattered and isolated precipitation
 674 and cloud features identifiable in observations. Model simulations at convective scales
 675 could help to improve the understanding of mesoscale dynamics for precipitation forma-
 676 tion over the TP.

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 681 NCEP/CPC Brightness temperatures can be downloaded from [https://disc.gsfc.nasa](https://disc.gsfc.nasa.gov/datasets/GPM_MERGIR_V1/summary)
 682 [.gov/datasets/GPM_MERGIR_V1/summary](https://disc.gsfc.nasa.gov/datasets/GPM_MERGIR_V1/summary) and GPM IMERG can be downloaded from
 683 [https://disc.gsfc.nasa.gov/datasets/GPM_3IMERGHH_06/summary?keywords=GPM](https://disc.gsfc.nasa.gov/datasets/GPM_3IMERGHH_06/summary?keywords=GPM%20IMERG%20v06%2030%20MIN)
 684 [%20IMERG%20v06%2030%20MIN](https://disc.gsfc.nasa.gov/datasets/GPM_3IMERGHH_06/summary?keywords=GPM%20IMERG%20v06%2030%20MIN). The code for the MCS tracking algorithm can be found
 685 at https://github.com/JuliaKukulies/mcs_tracking/tree/master/CTT/tracking.

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